# Epipolar resampling of linear pushbroom satellite imagery by a new epipolarity model ${ }^{\star}$ 

Mi Wang ${ }^{\text {a,* }}$, Fen Hu ${ }^{\text {b }}$, Jonathan $\mathrm{Li}^{\text {c }}$<br>${ }^{\text {a }}$ Liesmars, Wuhan University, No. 129, Luoyu Road Wuhan, Hubei 430079, China<br>${ }^{\mathrm{b}}$ Satellite Surveying and Mapping Application Center, State Bureau of Surveying and Mapping, Beijing, 100830, China<br>${ }^{\text {c }}$ University of Waterloo, Ontario, Canada

## A R T I CLE I N F O

## Article history:

Received 5 April 2010
Received in revised form 25 December 2010
Accepted 3 January 2011
Available online 5 February 2011

## Keywords:

Photogrammetry
Epipolar resampling
Linear pushbroom satellite imagery


#### Abstract

This paper presents a practical epipolarity model for high-resolution linear pushbroom satellite images acquired in either along-track or cross-track mode, based on the projection reference plane in object space. A new method for epipolar resampling of satellite stereo imagery based on this model is then developed. In this method, the pixel-to-pixel relationship between the original image and the generated epipolar image is established directly by the geometric sensor model. The approximate epipolar images are generated in a manner similar to digital image rectification. In addition, by arranging the approximate epipolar lines on the defined projection reference plane, a stereoscopic model with consistent ground sampling distance and parallel to the object space is thus available, which is more convenient for threedimensional measurement and interpretation. The results obtained from SPOT5, IKONOS, IRS-P5, and QuickBird stereo images indicate that the generated epipolar images all achieve high accuracy. Moreover, the vertical parallaxes at check points are at sub-pixel level, thus proving the feasibility, correctness, and applicability of the method.


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## 1. Introduction

Epipolar resampling aims to generate rectified images so that the stereo disparities are parallel to one of the image axes, thus providing advantageous conditions for automatic image matching, digital elevation model (DEM) generation, and stereo measurement (Cochran, 1995; Ono et al., 1999; Hashimoto, 2000; Bang et al., 2003; Lee et al., 2003; Morgan, 2004; Kornus et al., 2006). As is well known, epipolar resampling is a well-established procedure for images captured by a frame camera (Zhang, 1998; Torr, 2002), as the rigorous epipolar geometry can be easily defined and the epipolar lines are in a simple, straightforward linear style. However, as the leading payload of high-resolution imaging satellites, linear array pushbroom scanners are characterized by their particularly distinct imaging principle and their diverse, complex physical structures in comparison with conventional

[^0]frame cameras. That is, the image scenes are formed by stitching the one-dimensional (1D) scan lines captured as the sensor moves (Wolf and Dewitt, 2000). On the one hand, it is almost impossible to rigorously define the epipolar geometry of linear pushbroom satellite imagery, primarily because the determination of the accurate geometric relationship between object space and image space (i.e., sensor modeling) is difficult due to the adopted multicenter projection imaging mode. On the other hand, the resulting epipolar lines are actually not straight lines on the original image scenes, which brings complexity and difficulty to epipolar resampling. As a result, for satellite stereo imagery, the classical epipolar theory and methodology applied to single perspective frame-based imagery will no longer be feasible. Actually, epipolar resampling is the arrangement of epipolar lines based on a specific epipolarity model, so it is theoretically and practically significant to study the epipolarity model of linear pushbroom satellite stereo imagery.

In general, the existing epipolarity models can be categorized into two classes: (1) the models for finding epipolar curves in image space by image correspondence, and (2) the models for exploiting epipolar geometry based on a geometric sensor model. In the first category, the most often used model is the polynomial fit model (PFM), which was proposed and discussed to arrange the approximate epipolar curves of the SPOT cross-track stereo imagery in the late 1980s (Su, 2002). In the second category,
the Projection Trajectory based Epipolarity Model (PTEM) and the Parallel Projection transformation based Epipolarity Model (PPEM) are two typical examples. The Projection Trajectory (PT) based epipolar curve definition is most widely known and accepted for linear pushbroom satellite images (Alrousan et al., 1997), and can be extended to conventional frame-based images. Considerable effort has been expended on the PTEM and the major achievement has been the derivation of the mathematical form of the epipolar geometry from the rigorous sensor model based on collinearity equations (Kim, 2000; Bang et al., 2003; Zhao et al., 2008) and the simplified pushbroom sensor model (Lee and Park, 2002). Accordingly, the geometric properties of the theoretically nonstraight epipolar lines can thus be visibly analyzed, providing constructive instructions to the epipolarity related applications of linear pushbroom satellite images. In recent years, PPEM has been studied and applied to epipolar resampling of high-resolution linear pushbroom satellite imagery (e.g., IKONOS, QuickBird, etc.) (Morgan et al., 2004a,b, 2006; Jaehong et al., 2006). Due to the small field of view (FOV) of this type of spaceborne linear pushbroom scanner, the perspective projection sensor model of such satellite images is first transformed to a parallel projection sensor model (Fraser and Yamakawa, 2004), and then the epipolar geometry parameters can be determined by at least four pairs of conjugate points, according to the property that the parallel projection stereo imagery holds a unified epipolar line direction. Finally the epipolar images are generated in a manner similar to the process of digital image rectification. The feasibility of PPEM used for epipolar resampling towards automatic DEM generation from high-resolution satellite images has already been investigated (Morgan, 2004; Habib et al., 2004).

Although the existing models described above have been used to generate approximate epipolar images of satellite stereo imagery, technical limitations still exist when the generality, applicability, and simplicity of these models are taken into account. For example, epipolarity models established from either coordinate polynomial fitting or image projection trajectories are without exception in a non-linear style and the derivation of these models needs a certain number of qualified conjugate image points or image orientation parameters, which brings the great complexity to the epipolar resampling process. Moreover, the applicability of PPEM is actually not fully universal because the model needs to work well under the condition that the sensor's FOV is small, and the degree of approximation of the parallel projection model to the perspective projection model relates to topographic conditions and terrain relief, roll angle of the scanner, as well as some other factors. Moreover, compared with the rigorous perspective projection model, although it only requires a small number of parameters to describe the sensor model of parallel projection images holding straight epipolar lines parallel with each other, it needs a certain number of well-matched conjugate image points, along with ground control points (GCPs) and sometimes the roll angle of the imaging sensor to decide the parameters for perspective to parallel projection transformation, as well as epipolar resampling (Morgan et al., 2006; Jaehong et al., 2006). And all of these to a large extent restrict the degree of applicability of the PPEM. Currently, how to produce a qualified epipolar image directly from the provided high-accuracy orientation data of satellite stereo images, especially from the rational polynomial coefficients (RPC) of the rational function model (RFM), is yet an unsolved problem (Zhao et al., 2008) in the application field of high-resolution satellite imagery.

Under these circumstances, our previous work (Hu et al., 2008) initially proposed an epipolarity model of satellite stereo imagery based on the Virtual Horizontal Plane (VHP) in object space. The feasibility of our original idea was validated using only the VHPbased epipolar resampling method to generate the approximate
epipolar images of 5 m resolution panchromatic SPOT5-HRG stereo imagery. Therefore, the epipolarity model presented in our previous paper lacked depth and comprehensive analysis and validation. In addition, the concept of Virtual Horizontal Plane (VHP) is not very rigorous and understandable.

In this article, we aim to (1) investigate an epipolarity model for satellite stereo imagery based on the Projection Reference Plane (PRP) in object space, (2) develop a PRP-based method for epipolar resampling of linear pushbroom satellite images, (3) focus on universality of the method, by introducing various kinds of satellite images with different stereoscopic imaging modes and terrain conditions in the experiments, and finally (4) exploit some key factors that affect the accuracy of epipolar resampling results for optional technical routines and better performance of the epipolar resampling method. The newly proposed model needs no sensor model transformation as does PPEM, and needs no conjugate image points or ground control points (GCP) but just orientation data of satellite stereo images to decide the direction for epipolar rearrangement. Overall, the PRP-based method for epipolar resampling is more applicable and universal in practice.

The reminder of the paper is organized as follows. The next section introduces the principle and foundation of our PRP-based epipolarity model by recalling the geometric properties of epipolar curves of linear pushbroom satellite images, along with theoretical reasoning on practical image data. The following sections describe the workflow for generating the approximate epipolar images of pushbroom satellite stereo imagery based on the PRP in object space, followed by experimental results with accuracy assessment, further discussion, and finally the conclusions.

## 2. Principles

### 2.1. Sensor modeling for epipolar resampling

The effective epipolar resampling requires sound geo-locating accuracy of the geometric sensor models, so obtaining highquality orientation parameters of satellite stereo images is a prerequisite. Today, rigorous geometric sensor modeling of linear array pushbroom satellite imagery is still a challenging task. The calculation of image coordinates based on the rigorous geometric sensor model (i.e., collinearity equations) requires the best scan line search, which adds a considerable amount of computation for stereo mapping and the production of DEMs and digital orthoimage models (DOMs). Besides, vendors of highresolution commercial satellite images normally provide the endusers with the RPC parameters of the RFM instead of ephemeris and attitude parameters, as well as sensor information. It is demonstrated that the 0.01 pixel high fitting accuracy is totally possible with the 3rd-order form of the RFM, in comparison with the rigorous geometric sensor model (Yang, 2000). Therefore, the RFM can be a good replacement for the rigorous geometric sensor model in the epipolar resampling procedure with satellite stereo imagery. Sometimes the accuracy of the sensor-oriented RFM is not sufficient, and thus additional parameters for biascompensation (Fraser and Hanley, 2003; Fraser et al., 2006) are calculated using a certain number of GCPs or image tie points in order to ensure a fairly desirable level of geo-positioning accuracy.

### 2.2. Epipolar geometry of linear pushbroom imagery

The epipolar geometry of traditional frame-based stereo imagery is classically elaborated in many archives of photogrammetry, and it is well known that the parallel epipolar lines can be obtained by projecting the original non-parallel epipolar lines
on a pair of images parallel to the baseline. The generation of such a horizontal image pair, i.e., the epipolar images, is practically identical to digital image rectification. However, in a scene of linear pushbroom satellite imagery, each scan line has its own exterior orientation parameters at the exposure instant; it is difficult to establish the accurate geometric relationship between object space and image space due to the dynamic imaging mode and the highly correlated orientation parameters. According to the PT based epipolar curve definition, suppose $p$ is an image point on the left (right) image, if we project all points on the ray of $p$ to the right (left) image based on the image geometric sensor model, the locus of these projection points, named $e p$ here, is thus identified as the epipolar curve of $p$.

Suppose $p$ is located on an image scanline indexed with $i,\left(l_{p}, s_{p}\right)$ and $\left(0, y_{p}\right)$ are pixel coordinates and focal plane coordinates of $p$, respectively, $e$ is the size of sensor detector, $y_{0}$ is the translation parameter, then
$i=l_{p}$
$y_{p}=s_{p} \times e-y_{0}$.
Let $\left(X_{p}, Y_{p}, Z_{p}\right)$ be the coordinates of a certain object point on the ray of $p$; according to the collinearity conditions, there is

$$
\begin{align*}
\left(\begin{array}{c}
X_{p} \\
Y_{p} \\
Z_{p}
\end{array}\right) & =\left(\begin{array}{l}
X_{s_{s}} \\
Y_{S_{i}} \\
Z_{s_{i}}
\end{array}\right)+\lambda_{i} R_{i}\left(\begin{array}{l}
0 \\
y_{p} \\
-f
\end{array}\right) \\
& =\left(\begin{array}{l}
X_{s_{i}} \\
Y_{s_{i}} \\
Z_{s_{i}}
\end{array}\right)+\lambda_{i}\left(\begin{array}{lll}
r_{i 11} & r_{i 12} & r_{i 13} \\
r_{i 21} & r_{i 22} & r_{i 23} \\
r_{i 31} & r_{i 32} & r_{i 33}
\end{array}\right)\left(\begin{array}{l}
0 \\
y_{p} \\
-f
\end{array}\right) \tag{2}
\end{align*}
$$

where ( $X_{s_{i}}, Y_{s_{i}}, Z_{s_{i}}$ ) are coordinates of the projection center of the scanline indexed with $i, \lambda_{i}$ is the scale factor, $f$ is the focal length, $R_{i}$ is the rotation matrix of the scanline indexed with $i, r_{i m n}$ ( $m, n=$ $1,2,3$ ) are the elements of $R_{i}$.

Then, back-projecting ( $X_{p}, Y_{p}, Z_{p}$ ) to the right (left) image obtains the image point $q$. Similarly, suppose $q$ is located on the image scanline indexed with $j$, $\left(X_{s_{j}}, Y_{s_{j}}, Z_{s_{j}}\right)$ are coordinates of the projection center, $\left(l_{r}, s_{r}\right)$ and $\left(0, y_{r}\right)$ are pixel coordinates and focal plane coordinates of $q$ respectively, then
$j=l_{r}$
$y_{r}=s_{r} \times e-y_{0}$
and also

$$
\left(\begin{array}{l}
0  \tag{4}\\
y_{r} \\
-f
\end{array}\right)=\frac{1}{\lambda_{j}} R_{j}^{\prime}\left(\begin{array}{c}
X_{p}-X_{s_{j}} \\
Y_{p}-Y_{s_{j}} \\
Z_{p}-Z_{s_{j}}
\end{array}\right), \quad R_{j}^{\prime}=\left(\begin{array}{lll}
r_{j 11} & r_{j 21} & r_{j 31} \\
r_{j 12} & r_{j 22} & r_{j 32} \\
r_{j 13} & r_{j 23} & r_{j 33}
\end{array}\right)
$$

where $\lambda_{j}$ is the scale factor, $R_{j}^{\prime}$ is the rotation matrix of the scanline indexed with $j, r_{j m n}(m, n=1,2,3)$ are the elements of $R_{j}^{\prime}$.

When $\lambda_{i} \approx \lambda_{j}=\lambda$, combine Eqs. (2) and (4), to obtain

$$
\begin{align*}
\left(\begin{array}{l}
0 \\
y_{r} \\
-f
\end{array}\right) & =\frac{1}{\lambda_{j}} R_{j}^{\prime}\left(\left(\begin{array}{l}
X_{s_{i}}-X_{s_{j}} \\
Y_{s_{i}}-Y_{s_{j}} \\
Z_{s_{i}}-Z_{s_{j}}
\end{array}\right)+\lambda_{i} R_{i}\left(\begin{array}{l}
0 \\
y_{p} \\
-f
\end{array}\right)\right) \\
& \approx \frac{1}{\lambda} R_{j}^{\prime}\left(\begin{array}{l}
X_{s_{i}}-X_{s_{j}} \\
Y_{s_{i}}-Y_{s_{j}} \\
Z_{s_{i}}-Z_{s_{j}}
\end{array}\right)+R_{j}^{\prime} R_{i}\left(\begin{array}{l}
0 \\
y_{p} \\
-f
\end{array}\right) \tag{5}
\end{align*}
$$

Making certain transformations on Eq. (5) gives
$y_{r}=-f \frac{r_{j 11} A+r_{j 21} B+r_{j 31} C}{r_{j 13} A+r_{j 23} B+r_{j 33} C}$
here
$\lambda=\frac{-r_{j 11} \times X s_{i j}-r_{j 11} \times Y s_{i j}-r_{j 11} \times Z s_{i j}}{\left(r_{i 12} r_{j 11}+r_{i 22} r_{j 21}+r_{i 32} r_{j 31}\right) \times y_{p}-f\left(r_{i 13} r_{j 11}+r_{i 23} r_{j 21}+r_{i 33} r_{j 31}\right)}$
$A=X_{s_{i}}-X_{s_{j}}+\lambda\left(r_{i 12} \times y_{p}-r_{i 13} \times f\right)$
$B=Y_{s_{i}}-Y_{s_{j}}+\lambda\left(r_{i 22} \times y_{p}-r_{i 23} \times f\right)$
$C=Z_{s_{i}}-Z_{s_{j}}+\lambda\left(r_{i 32} \times y_{p}-r_{i 33} \times f\right)$.
For satellites moving with high stability, the attitude angles of a scene of linear pushbroom imagery are almost constant, and the position of the projection centers can be regarded as a linear function of exposure time or the scanline index, as shown in Eq. (7). Here, $m_{1}, m_{2}, m_{3}, n_{1}, n_{2}, n_{3}$ are the coefficients of oneorder polynomials.
$X_{s_{j}}=m_{1} \times l_{r}+n_{1}$
$Y_{s_{j}}=m_{2} \times l_{r}+n_{2}$
$Z_{s_{j}}=m_{3} \times l_{r}+n_{3}$.
Till then, the mathematical function of epipolar curve ep in a simplified form such as Eq. (8) can be derived from Eqs. (3), (6) and (7). Here, $k_{i}(i=1,2,3,4)$ are the coefficients that are definitely decided for image point $p$.
$k_{1} l_{r}+k_{2} s_{r}+k_{3} l_{r} s_{r}+k_{4}=0$.
Apparently, the derived mathematical models of epipolar curves are in a non-linear style and complicated with many orientation parameters included; the shape of the projection trajectory is actually not a straight line; besides, each point on an epipolar curve defined by PT has its specific conjugate epipolar curve, i.e., no conjugate epipolar curves exist theoretically.

Nevertheless, it is indicated that the epipolar curves of linear pushbroom satellite imagery hold two major geometric properties (Habib et al., 2004, 2005). One refers to the Straightness Approximation: although the epipolar curves are in a quasihyperbola shape at infinity (see Eq. (8)), they are almost straight lines over the extent of stereo image scenes; the other property is Local Conjugacy: although the conjugate epipolar curves do not exist in the traditional sense for linear pushbroom stereo images, they can still be considered to exist in a local area or close to the conjugate image points of stereo images by approximation. It is assumed that conjugate image points correspond with each other on these hypothetical conjugate epipolar curves.

### 2.3. PRP-based epipolarity model

As shown in Fig. 1, $0-X Y Z$ is the Local Vertical Coordinate system (LVCs), which is a left-handed coordinate system defined on the reference ellipsoid, with the origin point located near the central part of ground cover and with the $Y$ axis tangential to the reference ellipsoid and pointing in the due-north direction. The PRP is defined as the average elevation plane in object space with elevation value $H$. Suppose $p_{1}$ and $p_{2}$ are conjugate image points on the left and the right image, respectively, and they are intersected at ground point $P$. Derived from the PT based epipolar curve definition, $e p_{1}$ on the right image indicates the epipolar curve of point $p_{1}$, while $e p_{2}$ on the left image indicates the epipolar curve of point $p_{2}$. Inferred from the orbital stability of the satellite platform and the formerly depicted geometric properties of epipolar curves, $e p_{1}$ and $e p_{2}$ compose a pair of conjugate epipolar lines by approximation; moreover, if points on $e p_{1}$ and $e p_{2}$ are projected to the PRP, the projection trajectory will be very close to a straight line named $E D$. Besides, $E P$ decided by line $E D$ and point $P$ can be regarded as the approximate epipolar plane.


Fig. 1. The PRP-based epipolarity model of satellite stereo imagery. (a) Along-track stereo, and (b) cross-track stereo.

Moreover, it will not be difficult to discover that the PRP projection trajectories of all the conjugate epipolar curves can be approximated as a straight line set parallel with $E D$. To validate our theoretical considerations, coordinate calculation and visual inspection were applied to practical image data in this way: for $p_{1}$, select several evenly-distributed sample points from its epipolar curve (i.e., PT), and then calculate the coordinates of their projective points on PRP so as to get the locus; similarly, for $p_{2}$, the epipolar curve locus on PRP is also derived. As illustrated in Fig. 2, the horizontal axis and the vertical axis respectively plot the $X$ value and the $Y$ value of the projective points on PRP in the LVCs. Judging from the line equations fitted and the residuals from these
straight lines, it is obvious that for each pair of conjugate epipolar curves, the two loci almost coincide and with a high degree of straightness, while for different pairs of conjugate epipolar curves the loci are almost parallel.

That is to say, the approximate epipolar lines with parallelism can be generated by projecting the original satellite stereo imagery to an image pair parallel to the straight line $E D$, i.e., the equivalent line direction of loci on the PRP for epipolar curves, and thus the PRP-based epipolarity model can be proposed. Actually, the PRP of object space acts as a bridge to build up the pixel coordinate relationships between the original image and the generated epipolar image. Based on such an epipolarity model, it is possible to accomplish the epipolar resampling of satellite stereo imagery in a way that is identical to image digital rectification.

## 3. PRP-based epipolar resampling

This section describes a new method for epipolar image generation, according to the proposed PRP-based epipolarity model.

It is well known that the essence of the RFM is to associate the geographic coordinates of a ground point with the pixel coordinates of the corresponding image point, where coordinates of both ground point and image point are normalized between -1 and +1 so as to ensure the reliability of the parameter resolving process. The RFM has the following fundamental form:

$$
\begin{align*}
& x=\frac{\operatorname{Num}_{s}(U, V, H)}{\operatorname{Den}_{s}(U, V, H)} \\
& y=\frac{\operatorname{Num}_{L}(U, V, H)}{\operatorname{Den}_{L}(U, V, H)} \tag{9}
\end{align*}
$$

here

$$
\begin{aligned}
\operatorname{Num}_{L}(U, V, H)= & a_{1}+a_{2} V+a_{3} U+a_{4} H+a_{5} V U \\
& +a_{6} V H+a_{7} U H+a_{8} V^{2}+a_{9} U^{2} \\
& +a_{10} H^{2}+a_{11} U V H+a_{12} V^{3}+a_{13} V U^{2} \\
& +a_{14} V H^{2}+a_{15} V^{2} U+a_{16} U^{3}+a_{17} U H^{2} \\
& +a_{18} V^{2} H+a_{19} U^{2} H+a_{20} H^{3} \\
\operatorname{Den}_{L}(U, V, H)= & b_{1}+b_{2} V+b_{3} U+b_{4} H+b_{5} V U \\
& +b_{6} V H+b_{7} U H+b_{8} V^{2}+b_{9} U^{2} \\
& +b_{10} H^{2}+b_{11} U V H+b_{12} V^{3}+b_{13} V U^{2} \\
& +b_{14} V H^{2}+b_{15} V^{2} U+b_{16} U^{3}+b_{17} U H^{2} \\
& +b_{18} V^{2} H+b_{19} U^{2} H+b_{20} H^{3} \\
\text { Num }_{s}(U, V, H)= & c_{1}+c_{2} V+c_{3} U+c_{4} H+c_{5} V U+c_{6} V H \\
& +c_{7} U H+c_{8} V^{2}+c_{9} U^{2} \\
& +c_{10} H^{2}+c_{11} U V H+c_{12} V^{3}+c_{13} V U^{2} \\
& +c_{14} V H^{2}+c_{15} V^{2} U+c_{16} U^{3}+c_{17} U H^{2} \\
& +c_{18} V^{2} H+c_{19} U^{2} H+c_{20} H^{3}
\end{aligned}
$$



Fig. 2. PRP loci of four pairs of selected conjugate epipolar curves; the residuals from the fitted straight lines are $2 \mathrm{e}-5,4 \mathrm{e}-5,3 \mathrm{e}-4$ and $5 \mathrm{e}-5$ in meters respectively.

$$
\begin{aligned}
\operatorname{Den}_{s}(U, V, H)= & d_{1}+d_{2} V+d_{3} U+d_{4} H+d_{5} V U \\
& +d_{6} V H+d_{7} U H+d_{8} V^{2}+d_{9} U^{2} \\
& +d_{10} H^{2}+d_{11} U V H+d_{12} V^{3}+d_{13} V U^{2} \\
& +d_{14} V H^{2}+d_{15} V^{2} U+d_{16} U^{3}+d_{17} U H^{2} \\
& +d_{18} V^{2} H+d_{19} U^{2} H+d_{20} H^{3}
\end{aligned}
$$

where $a_{i}, b_{i}, c_{i}, d_{i}$ are the RPC parameters and $i=20 ;(U, V, H)$ and $(x, y)$ are the normalized coordinates of ground points and image points, respectively.

Suppose $(S, L)$ are the pixel coordinates of the image point, and (Lat, Lon, Height) are the geodetic coordinates of the corresponding object point, then
$U=\frac{\text { Lat }- \text { Lat_OFF }}{\text { Lat_SCALE }}, \quad V=\frac{\text { Lon }- \text { Lon_OFF }}{\text { Lon_SCALE }}$,
$H=\frac{H e i g h t ~-H_{-} O F F}{H_{-} S C A L E}$,
$x=\frac{S-S_{-} O F F}{S_{-} S C A L E}, \quad y=\frac{L-L_{-} O F F}{L_{-} S C A L E}$


Fig. 3. The decision of approximate line direction of the epipolar curve locus on PRP (strategy with no prior image matching process).

### 3.2. Determine PRP coverage of epipolar images

Based on Eqs. (11) and (15), coverage of the original left image on the PRP is calculated; similarly, based on Eqs. (13) and (15), PRP coverage of the original right image can also be calculated.

As shown in Fig. 1, AP denotes the PRP coverage of epipolar images, and is decided by the minimum rectangular region enveloping the entire PRP coverage of the original stereo imagery. Here, the horizontal boundary of AP is parallel to the line direction of $E D$, according to the principle of the PRP-based epipolarity model.

### 3.3. Epipolar resampling

Fig. 4 illustrates how to establish the pixel coordinate relation between the epipolar images and the original stereo imagery, which is the foundation of PRP-based epipolar resampling.

On the one hand, according to the principle of the PRP-based epipolarity model, AP is segmented into grid units with uniform size almost identical with the GSD (Ground Sampling Distance) of the original stereo imagery, and then the grid units of AP are associated with pixels of the left and right epipolar images. It is prescribed that the generated left and right epipolar images have the same PRP coverage as AP, so the left and right epipolar images are of the same size. As shown in Fig. 4, the corresponding relation between planarity coordinates of grid units in the LVCs and pixel coordinates of epipolar images can be established by rotation, translation and scaling transformations.

Suppose $O-X^{\prime} Y^{\prime}$ shares the same origin as that of $O-X Y Z$, and with its horizontal axis parallel to the line direction of $E D$. If $(X, Y)$ denote planarity coordinates of grid units in the LVCs, the corresponding coordinates in $0-X^{\prime} Y^{\prime}$ can be thus obtained by Eq. (17), where $R$ is the 2D rotation matrix derived from the line direction of $E D$ on PRP.
$\binom{X^{\prime}}{Y^{\prime}}=R \times\binom{ X}{Y}$.
Then, suppose ( $\mathrm{d} X^{\prime}, \mathrm{d} Y^{\prime}$ ) are the translation parameters from the origin of $O-X^{\prime} Y^{\prime}$ to the upper left corner of the AP area, $g$ is the grid size, $\left(x_{l}, y_{l}\right)$ and $\left(x_{r}, y_{r}\right)$ are pixel coordinates of the left and the right epipolar images, respectively, then
$\binom{x_{l} \times g}{y_{l} \times g}=\binom{X^{\prime}}{Y^{\prime}}+\binom{\mathrm{d} X^{\prime}}{\mathrm{d} Y^{\prime}}$
$\binom{x_{r} \times g}{y_{r} \times g}=\binom{X^{\prime}}{Y^{\prime}}+\binom{\mathrm{d} X^{\prime}}{\mathrm{d} Y^{\prime}}$.
Obviously, the above rotation, translation and scaling transformations are actually equivalent to Eq. (19), and the left and the


Fig. 4. Coordinate transformation relation of epipolar images and the original stereo imagery.
right epipolar images share the same set of 2D affine transformation parameters $a_{0}, a_{1}, a_{2}, b_{0}, b_{1}, b_{2}$.
$\left\{X=a_{0}+a_{1} \times x_{r}+a_{2} \times y_{r}\right.$
$\left\{Y=b_{0}+b_{1} \times x_{r}+b_{2} \times y_{r}\right.$
$\left\{\begin{array}{l}X=a_{0}+a_{1} \times x_{l}+a_{2} \times y_{l} \\ Y=b_{0}+b_{1} \times x_{l}+b_{2} \times y_{l} .\end{array}\right.$
On the other hand, for each of the grid units on PRP, the corresponding pixel position on the original left image can be calculated based on Eqs. (12) and (16); similarly, based on Eqs. (14) and (16), the corresponding pixel position on the original right image can also be computed.

Thus, the pixel coordinate relation between the epipolar images and the original stereo imagery is established, by taking the PRP as a bridge. Finally, an appropriate interpolation, e.g., bicubic, is usually performed to finish the entire epipolar resampling process.

## 4. Experiments and discussion

### 4.1. Dataset

To prove the theoretical feasibility of the PRP-based epipolarity model and the generality of the developed epipolar resampling approach, 5 m SPOT5-HRG, 1 m IKONOS, 2.5 m IRS-P5, and 0.6 m QuickBird panchromatic stereo images with different imaging constellations and terrain relief were used in this study. Table 1 summarizes the specifications of these four kinds of satellite stereo images.

### 4.2. Epipolar resampling and accuracy assessment

Before epipolar resampling, bias compensation is performed on the initially provided RPC orientation data in order to achieve a desirable sensor modeling accuracy. In the case of SPOT-HRG stereo imagery in which 21 GCPs are available, we selected half of them as control points to calculate the bias correction parameters
of the corresponding RPC orientation data, while the remaining half are used as check points to evaluate the result. In the case of IKONOS stereo imagery in which five control points are available, we used three of them as control points and two as check points to validate the accuracy after the RPC refinement. Since there were no GCPs available for the IRS-P5 and QuickBird images, bias compensation of the RPC orientation parameters was performed based on several pairs of conjugate image points to reach high relative orientation accuracy. The bias compensated RPC parameters of these images all achieved accuracy at the subpixel level.

With the refined RPC orientation parameters, the epipolar images of various satellite stereo images were produced. In addition, an ancillary file containing the 2D translation, rotation, and scaling parameters as well as RPC parameters was produced along with each pair of epipolar images. This ancillary file was used to record the pixel coordinate transformation relation between the epipolar images and the original stereo imagery for further utilizations. All the procedures used in the experiment are implemented in the Microsoft Visual C++ Integrated Development Environment on a desktop computer configured with the Microsoft XP Pro operating system.

The generated epipolar stereo models all achieved satisfactory stereoscopic visual effects. Fig. 5(a)-(d) give an overview of stereo anaglyphs of the generated epipolar images; in addition, Fig. 5(e) and (f) show stereo anaglyphs of a local zoomed region in IKONOS stereo imagery with different horizontal parallaxes, and it indicated that points on the roof (see white mark in 5(e)) and on the pavement close by (see white mark in 5(f)) all satisfy the epipolarity constraint.

In order to numerically validate the accuracy of the generated epipolar images, the Leica Photogrammetry Suite (LPS) module of ERDAS IMAGINE software was applied to extract 30 pairs of conjugate image points from each kind of satellite stereo imagery. These conjugate image points are reliable and almost fully distributed over the range of image space. Based on the pixel coordinate relation between the epipolar images and the original stereo imagery, the positions and thus vertical parallaxes

Table 1
Specification of the experimental image datasets.

| Stereo imagery | Acquisition date | Image size (pixels) | Terrain type | Mean elevation (m) | GSD (m) | Ground coverage (km) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| SPOT5-HRG | $\begin{aligned} & \text { 2002-08-15 } \\ & 2002-08-18 \end{aligned}$ | $\begin{aligned} & 12,000 \times 12,000 \\ & 12,000 \times 12,000 \end{aligned}$ | Mountainous, hilly area | 370 | 5 | $\begin{aligned} & 60 \times 60 \\ & 60 \times 60 \end{aligned}$ |
| IKONOS | $\begin{aligned} & 2002-09-07 \\ & 2002-09-07 \end{aligned}$ | $\begin{aligned} & 19,043 \times 7853 \\ & 19,043 \times 7648 \end{aligned}$ | Plain | 20 | 1 | $\begin{aligned} & 19 \times 8 \\ & 19 \times 8 \end{aligned}$ |
| IRS-P5 | $\begin{aligned} & 2005-01-16 \\ & 2005-01-16 \end{aligned}$ | $\begin{aligned} & 12,000 \times 12,000 \\ & 12,000 \times 12,000 \end{aligned}$ | Mountainous, hilly area | 100 | 2.5 | $\begin{aligned} & 30 \times 30 \\ & 30 \times 30 \end{aligned}$ |
| QuickBird | $\begin{aligned} & 2004-03-02 \\ & 2004-03-02 \end{aligned}$ | $\begin{aligned} & 35,170 \times 27,576 \\ & 35,170 \times 23,708 \end{aligned}$ | Mountainous, hilly area | 600 | 0.6 | $\begin{aligned} & 21 \times 17 \\ & 21 \times 14 \end{aligned}$ |



Fig. 5. Stereo anaglyphs of the generated epipolar images.
of these conjugate image points on the generated epipolar images are calculated.

The four line graphs in Fig. 6 respectively show the vertical parallaxes of conjugate image points on the generated SPOT5HRG, IKONOS, IRS-P5, and QuickBird epipolar images, where the horizontal axis shows the index of conjugate image points ordered by increasing parallax, while the vertical axis shows the value of vertical parallax.

Moreover, Table 2 gives the statistics where $Q_{\text {min }}, Q_{\text {Max }}, Q_{\text {Mean }}$ and $Q_{\text {RMS }}$ indicate minimum, maximum, mean and root mean square values of vertical parallaxes, respectively.

## Table 2

Vertical parallax statistics of conjugate image points on generated epipolar images (unit: pixel).

| Stereo imagery | $Q_{\text {Min }}$ | $Q_{\text {Max }}$ | $Q_{\text {Mean }}$ | $Q_{\text {RMS }}$ |
| :--- | :--- | :--- | :--- | :--- |
| SPOT5-HRG | -0.480 | 0.250 | -0.197 | 0.185 |
| IKONOS | -0.638 | 0.427 | 0.003 | 0.248 |
| IRS-P5 | -0.575 | 0.397 | -0.016 | 0.231 |
| QuickBird | -0.779 | 0.666 | -0.050 | 0.311 |

Judging from the results indicated by Fig. 6 and Table 2, the vertical parallaxes all reach sub-pixel level. It is clear that when the


Fig. 6. Vertical parallaxes of conjugate image points on the generated epipolar images.


Fig. 7. Deciding the approximate line direction of the epipolar curve locus on PRP (strategy with prior image matching process).
accuracy of image orientation data is satisfactory, the comparable rigorous satellite epipolar images with subtle vertical parallax can be generated by the proposed epipolar resampling method.

Through repetitive tests, we also find that when the value of parameter $h$ in Fig. 3 is replaced with different values (e.g., ranging from 20 to 400 m ), each time the vertical parallaxes of conjugate image points are almost the same, with variations in $Q_{\text {RMS }}$ being smaller than $1 \times 10^{-2}$ pixel. That is to say, the value of $h$ has little effect on the quality of epipolar resampling and thus can be chosen flexibly in practice.

Experiments also indicate that such an epipolar resampling method is also feasible with long-strip satellite stereo imagery. Generally speaking, the terrain relief has no effect on such a method, since the four kinds of satellite images used in our experiment actually cover various terrain types such as flat, hilly, and mountainous areas. Overall, it is the accuracy of the image orientation data that primarily affects the quality of the epipolar resampling result.

### 4.3. Further discussion

It is noted that the direction of $E D$ can be additionally determined by any pair of conjugate image points. As shown in Fig. $7, p_{1}$ and $p_{2}$ are the projection points of a pair of conjugate image points represented by $\boldsymbol{a}$ and $\boldsymbol{b}$. The direction of the line connecting $p_{1}$ and $p_{2}$ can be identified as the approximate arranging direction of epipolar lines on the PRP.

In this way, we successively use each pair of conjugate image points to determine the epipolar arranging direction, and then calculate the average and RMS values of the vertical parallaxes. The results listed in Table 3 indicate that it can serve as another strategy to decide the approximate line direction of the epipolar

Table 3
Vertical parallax statistics of conjugate image points on generated epipolar images (unit: pixel).

| Point no. | SPOT5-HRG |  | IKONOS |  | IRS-P5 |  | QuickBird |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $Q_{\text {Mean }}$ | $Q_{\text {RMS }}$ | $Q_{\text {Mean }}$ | $Q_{\text {RMS }}$ | $Q_{\text {Mean }}$ | $Q_{\text {RMS }}$ | $Q_{\text {Mean }}$ | $Q_{\text {RMS }}$ |
| 1 | -0.038 | 0.222 | -0.003 | 0.252 | 0.012 | 0.221 | 0.014 | 0.369 |
| 2 | 0.264 | 0.261 | 0.019 | 0.242 | 0.256 | 0.245 | 0.064 | 0.401 |
| 3 | 0.093 | 0.237 | 0.009 | 0.244 | -0.138 | 0.225 | 0.048 | 0.391 |
| 4 | 0.154 | 0.245 | -0.006 | 0.254 | -0.124 | 0.224 | -0.035 | 0.343 |
| 5 | 0.206 | 0.252 | 0.004 | 0.247 | 0.117 | 0.227 | -0.127 | 0.309 |
| 6 | -0.022 | 0.223 | 0.002 | 0.248 | 0.048 | 0.222 | -0.063 | 0.330 |
| 7 | 0.227 | 0.255 | 0.001 | 0.249 | -0.205 | 0.232 | -0.182 | 0.301 |
| 8 | 0.457 | 0.289 | -0.009 | 0.258 | -0.119 | 0.224 | -0.069 | 0.328 |
| 9 | 0.494 | 0.295 | -0.011 | 0.259 | 0.461 | 0.289 | -0.096 | 0.317 |
| 10 | -0.178 | 0.207 | -0.033 | 0.287 | 0.152 | 0.231 | -0.098 | 0.317 |
| 11 | -0.007 | 0.225 | 0.002 | 0.248 | -0.003 | 0.221 | -0.142 | 0.306 |
| 12 | 0.243 | 0.258 | 0.001 | 0.249 | -0.328 | 0.251 | -0.174 | 0.302 |
| 13 | 0.165 | 0.247 | 0.033 | 0.243 | 0.124 | 0.228 | 0.116 | 0.438 |
| 14 | -0.209 | 0.204 | 0.001 | 0.249 | -0.185 | 0.230 | 0.034 | 0.382 |
| 15 | 0.086 | 0.236 | 0.027 | 0.242 | -0.328 | 0.251 | -0.077 | 0.324 |
| 16 | 0.117 | 0.241 | 0.007 | 0.245 | 0.008 | 0.221 | 0.009 | 0.366 |
| 17 | -0.063 | 0.219 | -0.023 | 0.273 | 0.134 | 0.229 | -0.085 | 0.321 |
| 18 | -0.177 | 0.207 | -0.018 | 0.267 | -0.048 | 0.221 | -0.052 | 0.335 |
| 19 | -0.196 | 0.205 | 0.019 | 0.242 | 0.096 | 0.225 | -0.188 | 0.301 |
| 20 | -0.051 | 0.220 | -0.014 | 0.263 | 0.476 | 0.293 | 0.017 | 0.371 |
| 21 | 0.592 | 0.311 | -0.008 | 0.256 | 0.444 | 0.285 | -0.018 | 0.351 |
| 22 | -0.406 | 0.189 | 0.015 | 0.243 | -0.111 | 0.223 | 0.033 | 0.381 |
| 23 | -0.257 | 0.200 | 0.004 | 0.247 | 0.177 | 0.234 | -0.080 | 0.323 |
| 24 | -0.055 | 0.220 | 0.020 | 0.242 | -0.036 | 0.221 | -0.119 | 0.311 |
| 25 | 0.050 | 0.232 | 0.016 | 0.242 | 0.077 | 0.224 | 0.133 | 0.451 |
| 26 | -0.324 | 0.195 | 0.044 | 0.249 | -0.144 | 0.226 | -0.267 | 0.310 |
| 27 | -0.240 | 0.201 | -0.018 | 0.267 | -0.248 | 0.238 | -0.061 | 0.331 |
| 28 | 0.172 | 0.248 | 0.022 | 0.241 | -0.173 | 0.228 | -0.267 | 0.310 |
| 29 | -0.096 | 0.215 | -0.023 | 0.274 | 0.008 | 0.221 | -0.210 | 0.301 |
| 30 | -0.046 | 0.221 | -0.021 | 0.270 | -0.215 | 0.233 | -0.176 | 0.301 |

curve locus on the PRP, i.e., from just one pair of reliably-matched conjugate image points and the high-accuracy PRC parameters. However, when the pair of conjugate image points used here happens to intersect an object point on or quite near the PRP, unstable or incorrect results will be derived. Therefore, with regard to the procedure's simplicity and reliability, it is recommended that the former depicted strategy (see Fig. 3) without a prior image matching process be adopted in practice.

## 5. Conclusions

This paper extends our previous work on the epipolarity model for satellite stereo imagery based on the PRP in object space. We have presented a PRP-based epipolar resampling method for generating approximate epipolar images from linear pushbroom satellite images acquired either in the along-track or cross-track mode. SPOT5-HRG, IKONOS, IRS-P5, and QuickBird stereo images
have been used to quantitatively evaluate the proposed method. The experimental results demonstrate that our method works well on both along-track and cross-track linear pushbroom satellite stereo images with different stereoscopic viewing modes, spatial resolution, and terrain conditions. Our study also indicates that the accuracy of image orientation data generally determines the correctness of the epipolar resampling process.

In conclusion, the Approximated Straightness and Local Conjugacy properties of epipolar curves within the satellite image region provide important information on the proposition of our PRP-based epipolarity model. Also the successful use of this model for epipolar image generation again proves the practical applicability of these properties and provides experience for further research. Based on this model, the relationship of the pixel coordinates between the original image and the generated epipolar image can be directly established by geometric sensor model (or RFM) based computation, without prior sensor model transformation or image matching performances being needed. Moreover, by arranging the approximate epipolar lines on the defined PRP, the stereoscopic model that is parallel to the object space can be acquired with a consistent resolution. This helps to ensure an impressive and sound stereo-viewing effect. Overall, the PRP-based epipolar resampling method is simple, reliable, and applicable. So far, it has been successfully applied to the stereo mapping and automated DEM generation software modules of linear pushbroom satellite images developed at the State Key Laboratory for Information Engineering in Surveying, Mapping \& Remote Sensing (LIESMARS) at Wuhan University in China.

## Acknowledgements

This work was supported by the National High Technology Research and Development Program (2009AA12Z120) and the National Natural Science Foundation of China (40901209). It was also funded by Key Laboratory of Geo-informatics of State Bureau of Surveying and Mapping (No. 200902). The authors would like to thank the anonymous reviewers for their valuable comments.

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[^0]:    A A new method for generating approximate epipolar images from highresolution linear pushbroom satellite stereo images based on the projection reference plane in object space.

    * Corresponding author. Tel.: +86 27 68778969, +86 013971351686; fax: +86 27 68778969.

    E-mail addresses: wangmi@lmars.whu.edu.cn (M. Wang), hufen1984@163.com (F. Hu), junli@uwaterloo.ca (J. Li).

