# Quantum limits on estimating a waveform

Introduction. What's the problem?
 Standard quantum limit (SQL) for force detection. The right wrong story
 Beating the SQL. Three strategies

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## I. Introduction. What's the problem?



View from Cape Hauy Tasman Peninsula Tasmania

## Measuring a classical parameter

Phase shift in an (optical) interferometer Readout of anything that changes optical path lengths Michelson-Morley experiment Gravitational-wave detection Planck-scale, holographic uncertainties in positions

Torque on or free precession of a collection of spins Magnetometer Atomic clock

Force on a linear system Gravitational-wave detection Accelerometer Gravity gradiometer Electrometer Strain meter

### (Absurdly) high-precision interferometry for force sensing

#### Hanford, Washington





The LIGO Collaboration, Rep. Prog. Phys. 72, 076901 (2009).

### Laser Interferometer Gravitational Observatory (LIGO)



#### Livingston, Louisiana

## (Absurdly) high-precision interferometry for force sensing Initial LIGO

#### Hanford, Washington



 $\left(egin{array}{c} {
m differential} \ {
m strain} \ {
m sensitivity} \end{array}
ight)\simeq 10^{-21}$ 

differential displacement sensitivity  $\simeq 4 \times 10^{-18} \,\mathrm{m}$ 

from 40 Hz to 7,000 Hz.

### Laser Interferometer Gravitational Observatory (LIGO)



#### High-power, Fabry-Perot-cavity (multipass), powerrecycled interferometers

#### Livingston, Louisiana

### (Absurdly) high-precision interferometry for force sensing

#### Hanford, Washington



### **Advanced LIGO**

differential strain  $1 \simeq 3 \times 10^{-23}$ sensitivity

differential displacement )  $\simeq 10^{-19}\,{\rm m}$ sensitivity

from 10 Hz to 7,000 Hz.

### Laser Interferometer Gravitational Observatory (LIGO)



#### Livingston, Louisiana

**High-power**, Fabry-**Perot-cavity** (multipass), powerand signal-recycled, squeezed-light interferometers

# Opto, atomic, electro micromechanics



Atomic force microscope



T. Rocheleau, T. Ndukum, C. Macklin, J. B. Hertzberg, A. A. Clerk, and K. C. Schwab, Nature 463, 72 (2010).



# Opto, atomic, electro micromechanics





Drum microresonator

A. D. O'Connell *et al.,* Nature 464, 697 (2010).





M. Eichenfield, R. Camacho, J. Chan, K. J. Vahala, and O. Painter, Nature 459, 550 (2009).

A. Schliesser and T. J. Kippenberg, Advances in Atomic, Molecular, and Optical Physics, Vol. 58, (Academic Press, San Diego, 2010), p. 207.





# Mechanics for force sensing

T. J. Kippenberg and K. J. Vahala, Science 321, 172 (2008).



## Standard quantum limit (SQL)

#### Wideband detection of force *f* on free mass *m* LIGO interferometer

$$\Delta q \simeq \sqrt{\Delta q_0^2 + \frac{\Delta p_0^2 \tau^2}{m^2}} \ge \sqrt{\frac{2\tau \Delta q_0 \Delta p_0}{m}} \ge \sqrt{\frac{\hbar \tau}{m}} \equiv \Delta q_{\text{SQL}}$$

$$Back \text{ action}$$

$$\delta q \simeq \frac{f\tau^2}{2m} \implies f_{\text{SQL}} \equiv \frac{2m}{\tau^2} \Delta q_{\text{SQL}} = \sqrt{\frac{4\hbar m}{\tau^3}}$$

 $m\simeq 50\,{
m kg},\qquad \Delta 
u=1/ au\simeq 100\,{
m Hz}$ 

 $\implies \Delta q_{\rm SQL} \simeq 10^{-19} \, {\rm m}, \quad f_{\rm SQL} \simeq 100 \, {\rm fN}$ 

### Standard quantum limit (SQL) Narrowband, on-resonance detection of force f on oscillator of mass m and resonant frequency $\omega_0$ Nanoresonator

 $\Delta q_{\text{SQL}} \equiv \sqrt{\frac{\hbar}{2m\omega_0}} \quad \text{Back action?}$  $\delta q \simeq \frac{f\tau}{2m\omega_0} \implies f_{\text{SQL}} \equiv \frac{2m\omega_0}{\tau} \Delta q_{\text{SQL}} = \sqrt{\frac{2\hbar m\omega_0}{\tau^2}}$ 

 $m \simeq 10 \text{ pg}, \quad 1/\tau_0 = \omega_0/2\pi \simeq 10 \text{ MHz}, \quad Q \simeq 10^4 - 10^6$ 

 $\implies \Delta q_{\rm SQL} \simeq 10 \, {\rm fm}, \quad f_{\rm SQL} \simeq 100 \, {\rm fN} \times \frac{\tau_0}{\tau}$ 

 $\begin{pmatrix} \text{force between two} \\ \text{Bohr magnetons} \\ \text{separated by } r = 1 \text{ nm} \end{pmatrix} = \frac{\mu_0}{4\pi} \times \frac{\mu_B^2}{r^4} \simeq 10 \text{ aN} \\ \mu_B = e\hbar/2m_ec = e\lambda_c/4\pi \simeq e \times 0.2 \text{ pm} \end{cases}$ 

#### Wideband force f on free mass m

$$\Delta q_{\rm SQL} = \sqrt{\frac{\hbar\tau}{m}} \qquad f_{\rm SQL} = \sqrt{\frac{4\hbar m}{\tau^3}} = \Delta \nu \sqrt{4\hbar m (\Delta \nu)}$$

On-resonance force *f* on oscillator of mass *m* and resonant frequency  $\omega_0$ 

$$\Delta q_{\rm SQL} = \sqrt{\frac{\hbar}{2m\omega_0}} \qquad f_{\rm SQL} = \sqrt{\frac{2\hbar m\omega_0}{\tau^2}} = \Delta \nu \sqrt{2\hbar m\omega_0}$$

It's wrong.

It's not even the right wrong story.

The right wrong story. Waveform estimation.  $S_{SQL}(\omega) = \frac{\hbar}{|G(\omega)|} = \hbar m \sqrt{(\omega^2 - \omega_0^2)^2 + 4\beta^2 \omega^2}$ 

## II. Standard quantum limit (SQL) for force detection. The right wrong story



Oljeto Wash Southern Utah

# SQL for force detection









If shot noise dominates, squeeze the phase quadrature.



## Noise-power spectral densities

Zero-mean, time-stationary random process *u(t)* 

$$u(t) = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} u(\omega) e^{-i\omega t} \qquad u(\omega) = \int_{-\infty}^{\infty} dt \, u(t) e^{i\omega t}$$

Noise-power spectral density of *u* 

$$\langle u^2 \rangle = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} S_u(\omega)$$



### SQL for force detection $k = m\omega_0^2 \qquad m \qquad \text{Signal force } f(t) \text{ (waveform only on the second se$



## Langevin force



$$S_{L}(\omega) = 4m\beta\hbar\omega \left(\frac{1}{2} + \frac{1}{e^{\hbar\omega/kT} - 1}\right)$$

$$S_{SQL}(\omega) = \frac{\hbar}{|G(\omega)|} = \hbar m \sqrt{(\omega^{2} - \omega_{0}^{2})^{2} + 4\beta^{2}\omega^{2}}$$

$$Q = \frac{\omega_{0}}{2\beta} = 10^{3}$$

$$0.1$$

$$S_{SQL}(\omega) = \frac{10^{3}}{10^{4}}$$

$$S$$

## SQL for force detection

$$S_{\text{SQL}}(\omega) = \frac{\hbar}{|G(\omega)|} = \hbar m \sqrt{(\omega^2 - \omega_0^2)^2 + 4\beta^2 \omega^2}$$

# The right wrong story.

$$\frac{S_{\eta}(\omega)}{S_{\xi}(\omega)} = |G(\omega)|^2 \qquad S_{\eta}(\omega)S_{\xi}(\omega) = \hbar^2/4$$

In an opto-mechanical setting, achieving the SQL at a particular frequency requires squeezing at that frequency, and achieving the SQL over a wide bandwidth requires frequency-dependent squeezing.

## III. Beating the SQL. Three strategies



Truchas from East Pecos Baldy Sangre de Cristo Range Northern New Mexico

## Beating the SQL. Strategy 1



- 1. Couple parameter to observable *h*, and monitor observable *o* conjugate to *h*.
- 2. Arrange that *h* and *o* are *conserved* in the absence of the parameter interaction; *o* is the simplest sort of *quantum nondemolition* (QND) or *back-action-evading* (BAE) observable.
- 3. Give *o* as small an uncertainty as possible, thereby giving *h* as big an uncertainty as possible (back action).

Strategy 1. Monitor a quadrature component.

 $q = \operatorname{Re}[(X_1 + iX_2)e^{-i\omega_0 t}] = X_1 \cos \omega_0 t + X_2 \sin \omega_0 t$   $/\dots = \operatorname{Im}[(X_1 + iX_2)e^{-i\omega_0 t}] = X_1 \cos \omega_0 t + X_2 \sin \omega_0 t$ 

 $p/m\omega_0 = \text{Im}[(X_1 + iX_2)e^{-i\omega_0 t}] = -X_1 \sin \omega_0 t + X_2 \cos \omega_0 t$ 

#### **Downsides**

- 1. Detect only one quadrature of the force.
- 2. Mainly narrowband (no convenient free-mass version).
- 3. Need new kind of coupling to monitor oscillator.



W. G. Unruh, in Quantum Optics, Experimental Gravitation, and Measurement Theory, edited by P. Meystre and M. O. Scully (Plenum, 1983), p. 647; F. Ya. Khalili, PRD 81, 122002 (2010).

Beating the SQL. Strategy 2



Strategy 2. Squeeze the entire output noise by correlating the measurement and back-action noise.

$$y(\omega) = \underbrace{G(\omega)f(\omega) + G(\omega)\xi(\omega) + \eta(\omega)}_{= q(\omega)}$$

Squeeze this output noise by correlating  $\eta$  and  $\xi$ . Quantum mechanics requires that an orthogonal linear combination of  $\eta$  and  $G\xi$  become very noisy, thus making  $\eta$ ,  $\xi$ , and q very noisy.

# Quantum Cramér-Rao Bound (QCRB)

Single-parameter estimation: Bound on the error in estimating a classical parameter that is coupled to a quantum system in terms of the inverse of the quantum Fisher information.

Multi-parameter estimation: Bound on the covariance matrix in estimating a set of classical parameters that are coupled to a quantum system in terms of the inverse of a quantum Fisher-information matrix.

Waveform estimation: Bound on the continuous covariance matrix for estimating a continuous waveform that is coupled to a quantum system in terms of the inverse of a continuous, two-time quantum Fisher-information matrix.

Waveform QCRB.  
Spectral uncertainty principle  

$$S_{est}(\omega)\left(S_{\Delta q}(\omega) + \frac{\hbar^2}{4S_{\Delta f}(\omega)}\right) \geq \frac{\hbar^2}{4}$$
  
 $S_{\Delta q}(\omega) = |G(\omega)|^2 S_{\xi}(\omega)$ 

At frequencies where there is little prior information,

$$S_{\text{est}}(\omega) \geq \frac{\hbar^2}{4S_{\Delta q}(\omega)} = \frac{1}{|G(\omega)|^2} \frac{\hbar^2}{4S_{\xi}(\omega)} = \frac{S_{\eta}(\omega)}{|G(\omega)|^2}$$
  
Minimum-uncertainty noise

No hint of SQL—no back-action noise, only measurement noise—but can the bound be achieved?

Strategy 3. Quantum noise cancellation (QNC) using oscillator and negative-mass oscillator.

Beating the SQL. Strategy 3





#### Quantum noise cancellation M. Tsang an PRL 105,123

M. Tsang and C. M. Caves, PRL 105,123601 (2010).

Oscillator (q,p) and negative-mass oscillator (q',p')

$$Q = q + q' \implies P = (p + p')/2$$
  
$$\delta q = (q - q')/2 \implies \delta p = p - p'$$
  
Oscillator pairs

Back-action noise in q and  $q^\prime$  cancels in  $Q = q + q^\prime$  OR

Q = q + q' is a new BAE observable, which, rather than being conserved, acts just like oscillator position, responding to a force in the same way.

Paired sidebands about a carrier frequency Paired collective spins, polarized along opposite directions

> W. Wasilewski , K. Jensen, H. Krauter, J. J. Renema, M. V. Balbas, and E. S. Polzik, PRL 104, 133601 (2010).

### That's it, folks! Thanks for your attention.

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