

QFT for Cosmology, Achim Kempf, Lecture 13

Note Title

Recall: The free Klein-Gordon quantum field in a generic curved space-time must obey:

$$\hat{\phi}^+(x,t) = \hat{\phi}(x,t), \quad \hat{\pi}^+(x,t) = \hat{\pi}(x,t) \quad (\text{HC})$$

$$i\dot{\hat{\phi}}(x,t) = [\hat{\phi}(x,t), \hat{n}(t)], \quad i\dot{\hat{\pi}}(x,t) = [\hat{\pi}(x,t), \hat{n}(t)] \quad (\text{EoM})$$

which can be written in this form:

$$\left(\frac{1}{\sqrt{g}} \frac{\partial}{\partial x^\mu} \sqrt{g} g^{\mu\nu} \frac{\partial}{\partial x^\nu} + m^2 \right) \hat{\phi}(x,t) = 0, \quad \hat{\pi}(x,t) = \sqrt{g} g^{\mu\nu} \frac{\partial}{\partial x^\mu} \hat{\phi}(x,t) \quad (\text{EoM})$$

And: On all spacelike hypersurfaces, Σ , the CCRs must hold:

$$[\hat{\phi}(x,t), \hat{\phi}(y,t)] = 0, \quad [\hat{\pi}(x,t), \hat{\pi}(y,t)] = 0, \quad [\hat{\phi}(x,t), \hat{\pi}(y,t)] = i\hbar \delta^3(x-y) \quad (\text{CCR})$$

We want to show: The following ansatz for $\hat{\phi}(x,t)$ succeeds:

$$\hat{\phi}(x,t) := \sum_k u_k(x,t) a_k + u_k^*(x,t) a_k^*, \quad \text{with } [a_k, a_k^*] = \delta_{kk},$$

number-valued
solutions to K.G. eqn.

at least if the spacetime is globally hyperbolic.

So far we showed:

- The HC and EoM obeyed at all time.
- In a fixed coordinate system, CCRs are obeyed $\forall t$ if $\{u_k\}$ obey $\forall t$:

$$\sqrt{g} g^{\mu\nu} \sum_k \left(u_k(x,t) \frac{\partial}{\partial x^\mu} u_k^*(x,t) - u_k^*(x,t) \frac{\partial}{\partial x^\mu} u_k(x,t) \right) = i \delta^3(x-x') \quad (\text{W})$$

- Using Darboux's theorem, we showed that there exists a set of solutions $\{u_k\}$ so that (W) holds at some time to

Conservation of the CCRs? This is implied by the self-adjointness of \hat{H} :

□ As always in quantum theory, the time evolution

operator $\hat{U}(t, t_0) = T e^{i \int_{t_0}^t H(t') dt'}$ is unitary: $\hat{U}^\dagger = \hat{U}^{-1}$.

□ It allows one to express the time evolution of the observables, such as field operators, through:

$$\hat{\phi}(x, t) = \hat{U}(t, t_0) \hat{\phi}(x, t_0) \hat{U}^\dagger(t, t_0)$$

$$\hat{\pi}(x, t) = \hat{U}(t, t_0) \hat{\pi}(x, t_0) \hat{U}^\dagger(t, t_0)$$

□ Thus: $[\hat{\phi}(x, t), \hat{\pi}(x', t)] = [\hat{U} \hat{\phi}(x, t_0) \hat{U}^\dagger, \hat{U} \hat{\pi}(x', t_0) \hat{U}^\dagger]$

$$= \hat{U} [\hat{\phi}(x, t_0), \hat{\pi}(x', t_0)] \hat{U}^\dagger$$

$$= \hat{U} i \delta^3(x - x') \hat{U}^\dagger = i \delta^3(x - x')$$

Problem:

Is the quantization coordinate system independent?

Assume we solve the theory as above.

Now if we change coordinate system, and therefore the choices of $\{\Sigma\}$, would the CCRs still hold on every spacelike hypersurface Σ' ?

Proposition:

Yes: if CCRs hold in one coordinate system, then they hold in all: The CCRs keep holding when deforming a Σ to a Σ' .

Proof:

Rewrite the symplectic form (f, h) more abstractly:

Recall: $f, g \in V$ are solutions of KG eqn.

$$(f, h) := \int_{\Sigma} d\Sigma_r \sqrt{g} g^{mu} (f \partial_u h - h \partial_u f)$$

$$= \int_{\Sigma'} \underbrace{\tilde{j}}_{\text{a differential 3-form (Recall: Only 3-forms have 3-dim integrals)}}$$

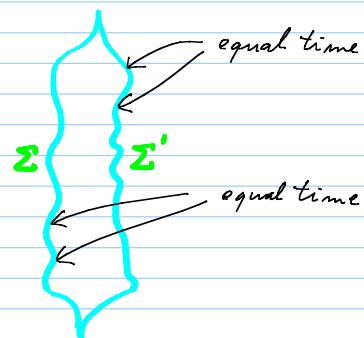
Here, we defined the contravariant vector field
 $j^\mu(x,t) := g^{\mu\nu} (f \partial_\nu h - h \partial_\nu f)$

and from it the differential 3-form:

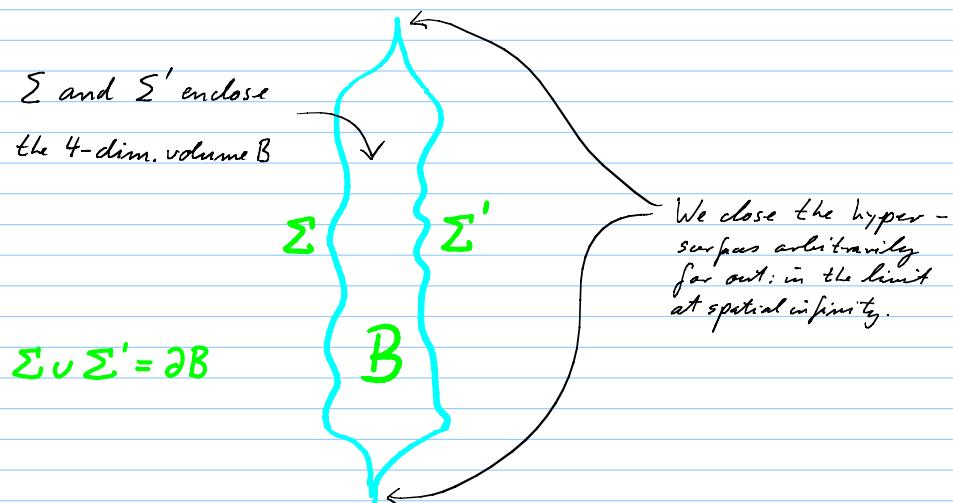
$$\tilde{j} := i_j \Omega$$

↑ inner derivation
 ↓
 3-form ↑ Volume 4-form $\sqrt{g} d^4x$

We need to show that the value of the symplectic form stays the same when deforming Σ :



□ Now integrate over both Σ and Σ' :



□ Use Stokes' theorem:

$$\int_{\partial B} \tilde{j} = \int_B d\tilde{j}$$

□ Notice:

If we can show $d\tilde{j} = 0$ we are done!

That's because then:

$$0 = \int_{\Sigma \cup \Sigma'} \tilde{j} = \int_{\Sigma} \tilde{j} + \int_{\Sigma'} \tilde{j} = - \int_{\Sigma} \tilde{j} + \int_{\Sigma'} \tilde{j}$$

Both j pointing
out of B , i.e one
to the future one to
the past.

Both j future
pointing.

$\Rightarrow \int_{\Sigma} \tilde{j}$ is indeed indep. of choice of Σ ,
if we can show $d\tilde{j} = 0$.

□ Indeed:

$$d\tilde{j} = d(i_j \Omega) = \text{div}_{\Omega} j = (\sqrt{g} j^{\mu})_{,\mu} d^4x$$

Here:

$$(\sqrt{g} j^{\mu})_{,\mu} = (\underbrace{\sqrt{g} g^{\mu\nu} (\cancel{f} \partial_{\nu} h - h \partial_{\nu} f)}_{\text{(in definition of } j\text{)}})_{,\mu}$$

$$= \cancel{\sqrt{g} g^{\mu\nu} \partial_{\nu} h} \partial_{\mu} f + f (\underbrace{\cancel{\sqrt{g} g^{\mu\nu} \partial_{\nu} h}}_{=-m^2 h \sqrt{g}})_{,\mu}$$

$$- \cancel{\sqrt{g} g^{\mu\nu} \partial_{\mu} h} \partial_{\nu} f - h (\underbrace{\cancel{\sqrt{g} g^{\mu\nu} \partial_{\nu} f}}_{=-m^2 f \sqrt{g}})_{,\mu}$$

Recall:

$(\square + m^2) \phi = 0$ reads:

$$\cancel{\sqrt{g} (\sqrt{g} g^{\mu\nu} \partial_{\nu} \phi)_{,\mu}} + m^2 \phi = 0$$

and the f and h are

solutions of the K. G. eqn!

$$= \cancel{\sqrt{g} (-f m^2 h + h m^2 f)} = 0 \quad \checkmark$$

→ We finally proved that, for globally hyperbolic spacetimes, there always exist mode functions $\{u_k(x,t)\}$ so that even ansatz for ϕ and Π also obeys the CCRs at all time and indeed $\forall \Sigma$:

$$\sqrt{g} g^{vv} \int \left(u_k(x,t) \frac{\partial}{\partial x^v} u_k^*(x',t) - u_k^*(x,t) \frac{\partial}{\partial x^v} u_k(x',t) \right) d^3k = i \delta^3(x-x') \quad (W)$$

Example: For Minkowski space, we had found this solution for the noninteracting Klein-Gordon field:

$$\hat{\phi}(x,t) = \int_{\mathbb{R}^3} \frac{1}{(2\pi)^{3/2}} \frac{1}{\sqrt{2\omega_k}} \left(a_k e^{-i\omega_k t + ikx} + a_k^* e^{i\omega_k t - ikx} \right) d^3k$$

We read off: $u_k(x,t) = \frac{1}{(2\pi)^{3/2}} \frac{1}{\sqrt{2\omega_k}} e^{-i\omega_k t + ikx}$

Now: Verify the CCR condition, (W):

□ Here: $\sqrt{g} = 1$ and $g^{vv} = \delta_{vv}$.

□ Thus, the LHS of Eq. (W) reads:

$$\int u_k(x,t) \frac{\partial}{\partial x^v} u_k^*(x',t) - u_k^*(x,t) \frac{\partial}{\partial x^v} u_k(x',t) d^3k$$

$$= \frac{1}{(2\pi)^3} \int \frac{1}{2\omega_k} \left[e^{-i\omega_k t + ikx} (i\omega_k) e^{i\omega_k t - ikx'} - e^{i\omega_k t - ikx} (-i\omega_k) e^{-i\omega_k t + ikx'} \right] d^3k$$

$$= \frac{1}{(2\pi)^3} \int \frac{2i\omega_k}{2\omega_k} e^{ik(x-x')} d^3k \stackrel{\text{Fourier}}{=} i \delta^3(x-x') \checkmark$$

Summary so far:

□ To solve the QFT of a free KG field on curved spacetime is to solve the **HC**, **EoM** and **CCRs**.

□ Make solution ansatz:

$$\hat{\psi}(x,t) := \sum_k u_k(x,t) a_k + u_k^*(x,t) a_k^* \quad (\text{A})$$

↑ or integral, e.g., if no IR cutoff

□ We showed that at least if spacetime is globally hyperbolic:

□ There exists a set of solutions of the KG eqns, $\{u_k\}$, so that ansatz (A) solves **HC**, **EoM** and **CCR** for all time.

Q: Does there exist only one such set of solutions?

A: No, there exist many other such sets of solutions: $\{\bar{u}_k\}, \{\tilde{u}_k\}, \dots$

How to see this non-uniqueness?

□ Recall symplectic form for $f, h \in V$:

$$(f, h) := \int_{\Sigma} d\Sigma_r \sqrt{g} g^{\mu\nu} (f \partial_{\mu} h - h \partial_{\mu} f)$$

□ Darboux: There exists a basis $\{v_m\}$ of V in which the form (\cdot, \cdot) reads:

$$\begin{pmatrix} 0 & & & \\ & \ddots & & \\ & & 0 & \\ 0 & & & \ddots \end{pmatrix}$$

□ From the v_m we constructed the $u_n := v_{2n} + i v_{2n+1}$

□ However: Darboux bases are not unique!

□ Example: 2-dim. solution subspace.

Assume: $v_1, v_2 \in V$ are a Darboux basis, so

so that $(,)$ reads $\begin{pmatrix} 0, 1 \\ -1, 0 \end{pmatrix}$.

Q: Can we change basis, $v_i = Q \bar{v}_i$, so that $(,)$ keeps that matrix form? Is there a matrix Q so that

$$Q^t \begin{pmatrix} 0, 1 \\ -1, 0 \end{pmatrix} Q = \begin{pmatrix} 0, 1 \\ -1, 0 \end{pmatrix} ?$$

$$\left(\begin{array}{l} \text{Since } (v, w) = v^t \begin{pmatrix} 0, 1 \\ -1, 0 \end{pmatrix} w \text{ we require} \\ v^t \begin{pmatrix} 0, 1 \\ -1, 0 \end{pmatrix} w = \bar{v}^t \begin{pmatrix} 0, 1 \\ -1, 0 \end{pmatrix} \bar{w} \quad \forall \bar{v}, \bar{w} \\ (\bar{B}\bar{v})^t \begin{pmatrix} 0, 1 \\ -1, 0 \end{pmatrix} (\bar{B}\bar{w}) = \bar{v}^t \begin{pmatrix} 0, 1 \\ -1, 0 \end{pmatrix} \bar{w} \quad \forall \bar{v}, \bar{w} \\ \text{i.e.: } \bar{B}^t \begin{pmatrix} 0, 1 \\ -1, 0 \end{pmatrix} \bar{B} = \begin{pmatrix} 0, 1 \\ -1, 0 \end{pmatrix} \end{array} \right) \rightarrow$$

A: Yes, any change of basis $Q = \begin{pmatrix} a & c \\ b & d \end{pmatrix}$

with $ad - bc = 1$ will do. (Exercise: check)

□ More generally: Also different boxes $\begin{pmatrix} 0, 1 \\ -1, 0 \end{pmatrix}$ can get mixed!

Exercise: □ Recall that with respect to the hermitian bi-linear form \langle , \rangle of Lecture 12, the mode functions $\{u_m\}$ obey:

$$\left. \begin{aligned} \langle u_n, u_m \rangle &= \delta_{nm} \\ \langle u_n^*, u_m^* \rangle &= -\delta_{nm} \\ \langle u_n, u_m^* \rangle &= 0 = \langle u_n^*, u_m \rangle \end{aligned} \right\} (*)$$

□ Now consider an invertible change of basis (in the space of complex-number-valued solutions of the K.B. eqn) to new mode functions:

$$\tilde{u}_m := \sum_n (A_{nm} u_m + B_{nm} u_m^*)$$

□ Show that for the $\{\tilde{u}_m\}$ to qualify as mode functions, i.e., for them to obey $(*)$, i.e., $\langle \tilde{u}_n, \tilde{u}_m \rangle = \delta_{nm}$ etc, A, B must obey:

$$A^* A - B^* B = 1 \text{ and } A^* B - B^* A^* = 0$$

Non-uniqueness of the solution (A):

□ Clearly, this means that there are infinitely many solutions of the form of (A) to HC, EoM and CCRs:

$$\hat{a}(x,t) := \sum_n u_n(x,t) a_n + u_n^*(x,t) a_n^*$$

$$\hat{a}(x,t) := \sum_n \bar{u}_n(x,t) \bar{a}_n + \bar{u}_n^*(x,t) \bar{a}_n^*$$

$$\hat{a}(x,t) := \sum_n \bar{\bar{u}}_n(x,t) \bar{\bar{a}}_n + \bar{\bar{u}}_n^*(x,t) \bar{\bar{a}}_n^* \quad \text{etc., etc...}$$

□ Correspondingly, we obtain different Fock bases:

Either: $a_n |0\rangle = 0 \quad |n_n\rangle := \frac{1}{\sqrt{n!}} (a_n^*)^n |0\rangle$

Or: $\bar{a}_n |0\rangle = 0 \quad |\bar{n}_n\rangle := \frac{1}{\sqrt{n!}} (\bar{a}_n^*)^n |0\rangle \quad \text{etc., etc...}$

□ Q: Do these solutions of the QFT

○ describe different physics, or

○ do they differ by a mere change of basis in Fock space and so describe the same physics?

□ A: It depends!

1) Assume first we can impose IR and UV cutoffs with negligible consequences

* This means we truncate to a finite (though large) number of independent mode oscillators, a_n, a_n^* .

* Then, the following theorem implies that all solutions to HC, EoM, CCR differ merely by a change of basis:

Theorem (Stone and von Neumann):

* Assume in a Hilbert space \mathcal{H} , the operators \hat{x}_i, \hat{p}_j obey:

$$[\hat{x}_i, \hat{p}_j] = i \delta_{ij} \quad [\hat{x}_i, \hat{x}_j] = 0 = [\hat{p}_i, \hat{p}_j] \quad i, j \in \{1, \dots, N\}$$

* Assume that in a Hilbert space $\tilde{\mathcal{H}}$ other operators \tilde{x}_i, \tilde{p}_j also obey:

$$[\tilde{x}_i, \tilde{p}_j] = i \delta_{ij} \quad [\tilde{x}_i, \tilde{x}_j] = 0 = [\tilde{p}_i, \tilde{p}_j] \quad i, j \in \{1, \dots, N\}$$

* Assume that the representations are irreducible (i.e., no invariant subspace).

We can assume that $\mathcal{H}' = \mathcal{H}$ because all separable Hilbert spaces (the usual: with countable bases) are unitarily equivalent.

Then: there exists a unitary operator \hat{U} so that:

$$\tilde{x}_i = U \hat{x}_i U^* \quad \tilde{p}_i = U \hat{p}_i U^* \quad (\text{i.e., a change of basis})$$

* Remark: Strictly speaking, there can be pathological cases.

The pathological cases can be avoided by requiring representations of the CCRs of the (bounded and therefore better behaved) operators:

$$e^{i\hat{p}_i \hat{x}_i}, e^{i\hat{x}_i \hat{p}_i}$$

* Application to QM and to UV&IR regularized $Q + T$:

$$\text{Consider } \hat{X}_n := \frac{1}{\sqrt{2}} (a_n + a_n^+), \quad \hat{P}_n := \frac{-i}{\sqrt{2}} (a_n - a_n^+)$$

$$\text{and } \bar{X}_n := \frac{1}{\sqrt{2}} (\bar{a}_n + \bar{a}_n^+), \quad \bar{P}_n := \frac{-i}{\sqrt{2}} (\bar{a}_n - \bar{a}_n^+) \text{ etc.}$$

The theorem of Stone & v. Neumann implies that

$$a_n = \hat{U} \bar{a}_n \hat{U}^* \quad \text{with } \hat{U} \text{ unitary.}$$

\Rightarrow All solutions are the same up to a mere change of basis.

2.) Consider now the possibility that we cannot truncate to a finite number of degrees of freedom.

Q: When would this happen?

A: E.g., phase transitions formally need systems with an infinite number of degrees of freedom.

Then: The QFT can have unitarily non-equivalent solutions, that differ physically: different "phases".

Underlying math of non-equivalent representations?

Assume $\langle a|b\rangle = d$ with $0 < d < 1$, i.e., not \perp

Then $(\underbrace{ka|ka|ka|\dots|a|}_{N})(\overbrace{|b\rangle|b\rangle|b\rangle|\dots|b\rangle}^N) = d^N$, i.e., not \perp

But for $N=\infty$ have $|a\rangle|a\rangle|\dots|a\rangle \perp |b\rangle|b\rangle|\dots|b\rangle$, so that then can no longer use $|a\rangle|a\rangle|\dots|a\rangle$ to help linearly combine, e.g., $|b\rangle|b\rangle|\dots|b\rangle$.

From now on: We will assume IR & UV cutoffs are possible and that Stone v. Neumann therefore applies.

Therefore:

□ No matter which set of suitable mode functions

$$\{u_n(x,t)\} \text{ or } \{\bar{u}_n(x,t)\} \text{ or } \{\bar{\bar{u}}_n(x,t)\}, \dots$$

we choose, we obtain the same solution

$$\hat{a}(x,t) = \sum_k u_k(x,t) a_k + u_k^*(x,t) a_k^*$$

$$= \sum_k \bar{u}_k(x,t) \bar{a}_k + \bar{u}_k^*(x,t) \bar{a}_k^*$$

$$= \sum_k \bar{\bar{u}}_k(x,t) \bar{\bar{a}}_k + \bar{\bar{u}}_k^*(x,t) \bar{\bar{a}}_k^* = \dots$$

with their Fock bases being different bases in the same Hilbert space.

□ For example, using the $\{u_n\}$, we are led to span the Hilbert space \mathcal{H} using this ON basis:

$|0\rangle$ where $a_k|0\rangle = 0 \quad \forall k$

$a_n^+|0\rangle, \frac{1}{\sqrt{n!}}(a_n^+)^n|0\rangle$

$\frac{1}{\sqrt{\dots}}(a_{k_1}^+)^{k_1}(a_{k_2}^+)^{k_2}\dots(a_{k_n}^+)^{k_n}|0\rangle, \text{etc...}$

□ Or, using other mode functions, say $\{\bar{u}_n\}$, we may span the same Hilbert space, \mathcal{H} , using this ON basis:

$|\bar{0}\rangle$ where $\bar{a}_k|\bar{0}\rangle = 0 \quad \forall k$

$\bar{a}_n^+|\bar{0}\rangle, \frac{1}{\sqrt{n!}}(\bar{a}_n^+)^n|\bar{0}\rangle$

$\frac{1}{\sqrt{\dots}}(\bar{a}_{k_1}^+)^{k_1}(\bar{a}_{k_2}^+)^{k_2}\dots(\bar{a}_{k_n}^+)^{k_n}|\bar{0}\rangle, \text{etc...}$

Does the choice of mode functions matter?

□ In principle, it does not:

Any state of the system, say $|\Psi\rangle$, can be expanded in each basis.

□ In practice, however:

It is convenient, whenever we know which state is the no-particle (i.e., vacuum) state, say $|\Omega\rangle$, to choose the mode functions $\{u_n\}$ such that the corresponding $|0\rangle$ is $|\Omega\rangle$, i.e., such that

$|0\rangle = |\Omega\rangle, \text{ i.e., such that } a_k|0\rangle = 0$

Then, conveniently, states like $\frac{1}{\sqrt{n!}}(a_n^+)^n|0\rangle$ are the multi-particle states.

Outlook: (only a rough sketch)

- Say we know the system's state, $|q\rangle$, is the vacuum initially.
- \rightsquigarrow We choose $\{a_m\}$ appropriately, so that $|0\rangle_{in} = |q\rangle$.
- After some evolution (e.g. the universe expands) the vacuum state may be a different state, say $|x\rangle$.
- \rightsquigarrow We choose $\{\bar{a}_m\}$ appropriately, so that $|\bar{0}\rangle_{out} = |x\rangle$
- At late times, since we work in the Heisenberg picture, the system is still in the state $|0_{in}\rangle$, but this is then an excited state!
 - \rightsquigarrow Description of particle production due to cosmic expansion.
- Recall: We had an analogous situation with driven harmonic oscillators!