

QFT for Cosmology, Achim Kempf, Lecture 15

Note Title

Solving the quantized K.G. eqns. on FRW spacetimes

Recall:

1.) We obtain the solution $\hat{\phi}(x,t)$ through the ansatz:

$$\hat{\phi}(x,t) = \sum_k u_k(x,t) a_k + u_k^*(x,t) a_k^* \quad (\star)$$

* if we use operators a_k obeying $[a_k, a_{k'}^\dagger] = \delta^3(k-k')$ and

* if we find classical solutions $\{u_k(x,t)\}$ of the K.G. eqn., called mode functions, which obey:

$$\sqrt{g} g^{xx} \sum_k \left(u_k(x,t) \frac{\partial}{\partial x^i} u_k^*(x,t) - u_k^*(x,t) \frac{\partial}{\partial x^i} u_k(x,t) \right) = i \delta^3(\vec{x} - \vec{x}') \quad (\dagger)$$

2.) Then, we can use the $\{a_k\}$ to build a convenient basis in the Hilbert space:

□ Namely: $|0\rangle$ is the vector obeying $a_k |0\rangle = 0$

□ The other basis vectors are:

$$a_k^+ |0\rangle, \dots \frac{1}{\sqrt{m!}} (a_k^+)^m |0\rangle, \dots a_k^+ a_{k'}^+ |0\rangle, \dots$$

$$\frac{1}{\sqrt{m_1!} \dots \sqrt{m_m!}} (a_{k_1}^+)^{m_1} \dots (a_{k_m}^+)^{m_m} |0\rangle, \dots \text{etc.}$$

3.) Choosing a different set of classical solutions $\{\tilde{u}_k(x,t)\}$ which obey (G) yields the same $\hat{\phi}(x,t)$, namely

$$\hat{\phi}(x,t) = \sum_k \tilde{u}_k(x,t) \tilde{a}_k + \tilde{u}_k^*(x,t) \tilde{a}_k^*$$

but the basis of vectors $|0\rangle, \tilde{a}_k^+ |0\rangle, \tilde{a}_k^+ \tilde{a}_{k'}^+ |0\rangle \dots$ is a different basis. Recall: Stone von Neumann theorem.

Application to FRW spacetime

For convenience (namely, to avoid a "friction"-type term) we aim to solve not for $\dot{\phi}(x, t)$ directly, but instead for:

$$\hat{x}(\eta, x) := a(\eta) \dot{\phi}(\eta, x)$$

In terms of $\hat{x}(\eta, x)$ the quantum K.G. eqn. reads:

$$\hat{x}''(\eta, x) - \Delta \hat{x}(\eta, x) + \left(m^2 a^2(\eta) - \frac{a''(\eta)}{a(\eta)} \right) \hat{x}(\eta, x) = 0$$

Note: This is a partial differential equation because both time and space derivatives occur.

Observation: The derivatives $\frac{\partial}{\partial x_i}$ become multiplication operators $i k_i$ under spatial Fourier transform.

Plan: Before trying to solve it, use Fourier to transform the K.G. eqn. from a partial DE into a more manageable set of ordinary DEs.

Define: $\hat{x}_k(\eta) := \int \frac{1}{(2\pi)^{3/2}} \hat{x}(\eta, x) e^{-ikx} d^3x$

$$\text{i.e.: } \hat{x}(\eta, x) = \int \frac{1}{(2\pi)^{3/2}} \hat{x}_k(\eta) e^{ikx} d^3k$$

Analogously:

$$\hat{\Pi}_k^{(xc)}(\eta) := \int \frac{1}{(2\pi)^{3/2}} \hat{\Pi}^{(xc)}(x, \eta) e^{-ikx} d^3x$$

□ Thus, in terms of $\hat{x}_k(\eta)$, the K.G. eqn. reads:

$$\hat{x}_k''(\eta) + \left(k^2 + m^2 a^2(\eta) - \frac{a''(\eta)}{a(\eta)} \right) \hat{x}_k(\eta) = 0 \quad (\text{EoM})$$

⇒ for each *comoving Fourier mode* k the K.G. eqn. is the eqn. of a harmonic oscillator with time-dependent frequency

$$\hat{x}_k''(\eta) + \omega_k^2(\eta) \hat{x}_k(\eta) = 0$$

$$\text{with: } \omega_k(\eta) := \sqrt{k^2 + m^2 a^2(\eta) - \frac{a''(\eta)}{a(\eta)}}$$

Remark: It is not surprising that the frequency of each comoving mode is changing because its physical wavelength is changing too.

In extreme cases:

The frequency $\omega_k(\eta)$ may become imaginary, namely if $a''(\eta)$ is large enough, i.e., if the expansion is rapid enough. Note that the discriminant also depends on k , i.e., some modes may have imaginary frequencies while others don't.

□ Exercise:

* Show that $\hat{x}_k^+(\eta) = \hat{x}_{-k}^-(\eta)$, $\hat{\pi}_k^{(\omega)+}(\eta) = \hat{\pi}_{-k}^{(\omega)-}(\eta)$ (HC)

* Show that

$$[\hat{x}_k(\eta), \hat{\pi}_{k'}(\eta)] = i \delta^3(k+k') \quad (\text{CCR})$$

$$\text{i.e. } [\hat{x}_k(\eta), \hat{\pi}_{k'}^+(\eta)] = i \delta^3(k-k')$$

□ In order to solve EoM, HG, CCR for $x_k(\gamma)$, we make this ansatz:

$$\hat{x}_k(\gamma) := \frac{1}{\sqrt{2}} \left(v_k^*(\gamma) a_k + v_k(\gamma) a_k^* \right) \quad (A)$$

convenient later

□ Exercise: Express the mode functions $u_k(\gamma, x)$ of (★) in terms of the functions $v_k(\gamma)$.

□ Proposition: The ansatz (A)

1.) solves the hermiticity condition (HG) by construction.

2.) solves the (EoM), if the $v_k(\gamma)$ each solve (EoM) as (complex!) number-valued functions:

Note: The equation depends only on $|k|$, not on the direction of k . Thus if $v_k(\gamma)$ is a solution for one k then it is solution for all k' with $|k'| = |k|$. \Rightarrow We can and will choose $v_k(\gamma) = v_{|k|}(\gamma)$

$$v_k''(\gamma) + \left(k^2 + m^2 a^2(\gamma) - \frac{a''(\gamma)}{a(\gamma)} \right) v_k(\gamma) = 0 \quad (H)$$

3.) the commutation relations (CCR) if the v_k are chosen such that they also obey:

$$v_k'(\gamma) v_k^*(\gamma) - v_k(\gamma) v_k^*(\gamma)' = 2i \quad (W)$$

□ Exercise:

a.) Prove the proposition.

b.) Assume that $v_k(\gamma)$ is any solution of (EoM).

Show that if (W) holds at one time then it holds at all time.

(Note: The LHS of (W) is the "Wronskian" of the ODE)

Conclusion:

In order to obtain the solution $\hat{\phi}(q, x)$, we do:

A) Find for each mode $k \in \mathbb{R}^3$ a solution $v_k(q)$ to (M), i.e., a solution to the classical harmonic oscillator with time-dependent frequency.

B) Make sure $v_k(q)$ always (W), if need be by multiplying with a constant. (Recall exercise b))

C) Build a basis in the Hilbert space:

$$a_k |0\rangle = 0, \quad a_k^\dagger |0\rangle, \quad a_k^\dagger a_k^\dagger |0\rangle, \text{ etc...}$$

Choice of mode solutions $\{v_k(q)\}$

□ For each choice, say $\{v_k(q)\}_{k \in \mathbb{R}^3}$ or $\{\tilde{v}_k(q)\}_{k \in \mathbb{R}^3}$ we obtain the same $\hat{\phi}(x, t)$ but the bases

$$|0\rangle, a_k^\dagger |0\rangle, a_k^\dagger a_k^\dagger |0\rangle, \dots$$

and

$$|\tilde{0}\rangle, \tilde{a}_k^\dagger |\tilde{0}\rangle, \tilde{a}_k^\dagger \tilde{a}_k^\dagger |\tilde{0}\rangle, \dots$$

will of course be different.

□ We will often find it convenient to use the basis $|0\rangle, a_k^\dagger |0\rangle, a_k^\dagger a_k^\dagger |0\rangle, \dots$ that comes with one set of mode functions $\{v_k(q)\}$ at one time (say initially) and then the basis $|\tilde{0}\rangle, \tilde{a}_k^\dagger |\tilde{0}\rangle, \tilde{a}_k^\dagger \tilde{a}_k^\dagger |\tilde{0}\rangle, \dots$ of some $\{\tilde{v}_k(q)\}$ later.

Why?

□ In the Heisenberg picture, the system's state vector is always the same Hilbert space vector.

□ But the observables evolve in time!

⇒ The meanings of all Hilbert space vectors change over time

→ We may, e.g., choose a set $\{v_k\}$ whose vector $|0\rangle$ happens to be the vacuum state at one time and we may choose another set $\{\tilde{v}_k\}$ whose vector $|\tilde{0}\rangle$ happens to be the vacuum state at another time.

□ How many possible choices of

$$\{v_k(\gamma)\}_{k \in \mathbb{N}^3}, \{\tilde{v}_k(\gamma)\}_{k \in \mathbb{N}^3}, \{\hat{v}_k(\gamma)\}_{k \in \mathbb{N}^3}, \dots$$

do exist?

□ Let us consider each mode, k , separately:

□ The solution space of (M), for fixed k ,

$$v_k''(\gamma) + \left(k^2 + m^2 a^2(\gamma) - \frac{a''(\gamma)}{a(\gamma)}\right) v_k(\gamma) = 0$$

has of course 2 complex dimensions.

[Note: Every solution obeying (M) must be complex-valued. Why?]

□ If v_k is a complex-valued solution, then

$$v_k \text{ and } v_k^\star$$

form a basis in the solution space.



Every solution, \tilde{v}_k , is a linear combination of v_k, v_k^* , i.e., there must exist $\alpha, \beta \in \mathbb{C}$, so that:

$$\tilde{v}_k(\gamma) = \alpha_k v_k(\gamma) + \beta_k v_k^*(\gamma)$$

□ The actual dimensionality is 3!

The solution space thus has 4 real dimensions, but one real dimension is lost because the solutions v_k, \tilde{v}_k , etc. must also obey (W), i.e.:

$$v_k'(\gamma) v_k^*(\gamma) - v_k(\gamma) v_k^{\prime*}(\gamma)' = 2i \quad (W)$$

(i.e. $\text{Im}(v' v^*) = 1$, which is only one real equation)

□ Proposition:

Assume v_k obeys (W). Then, \tilde{v}_k defined through

$$\tilde{v}_k(\gamma) = \alpha_k v_k(\gamma) + \beta_k v_k^*(\gamma) \quad (B)$$

also obeys (W), iff the coefficients α_k, β_k obey:

$$|\alpha_k|^2 - |\beta_k|^2 = 1$$

□ Proof: Exercise

$$\begin{aligned} \Rightarrow \text{we easily obtain: } \tilde{v}_k(\gamma) &= \frac{1}{\sqrt{2}} (v_k^*(\gamma) a_k + v_k(\gamma) a_{-k}^*) \\ &= \frac{1}{\sqrt{2}} (\tilde{v}_k^*(\gamma) \tilde{a}_k + \tilde{v}_k(\gamma) \tilde{a}_{-k}^*) \quad \left. \right\} (P) \\ &= \dots \end{aligned}$$

□ Terminology: Such a transformation from one choice $\{v_n\}, a_n$ and corresponding basis

$$|0\rangle, a_k^+ |0\rangle, a_k^+ a_n^+ |0\rangle, \dots$$

to some $\{\tilde{v}_n\}, \tilde{a}_n$ and their basis

$$|\tilde{0}\rangle, \tilde{a}_k^+ |\tilde{0}\rangle, \tilde{a}_k^+ \tilde{a}_n^+ |\tilde{0}\rangle, \dots$$

is called a "Bogoliubov transformation".

Strategy: We have two tasks now:

* Make Bogoliubov Hilbert basis transforms explicit.

(E.g. so that $|0\rangle$ is, at least at one time, the vacuum.) * Find out when which choice of $\{v_n\}$ is convenient.

Bogoliubov transformations of Hilbert bases

□ How can we express the basis vectors

$$|\tilde{0}\rangle, \tilde{a}_k^+ |\tilde{0}\rangle, \frac{1}{\sqrt{a!}} \tilde{a}_k^{+2} |\tilde{0}\rangle, \dots, \tilde{a}_k^+ \tilde{a}_n^+ |\tilde{0}\rangle, \dots$$

as linear combinations of the basis

$$|0\rangle, a_n^+ |0\rangle, \frac{1}{\sqrt{a!}} a_n^{+2} |0\rangle, \dots, a_k^+ a_n^+ |0\rangle, \dots ?$$

□ Proposition: Equations (B) & (P) yield: $a_n = d_n^* \tilde{a}_n + f_n \tilde{a}_{-n}^*$

Proof: Exercise.

□ Now we observe that $a_n |0\rangle = 0$ becomes:

$$(d_n^* \tilde{a}_n + f_n \tilde{a}_{-n}^*) |0\rangle = 0$$

□ Try to solve for $|1\rangle$ using ansatz: $|1\rangle := \left(\prod_k f_k(\hat{a}_k^+, \hat{a}_{-k}^+) \right) |1\rangle$

□ Proposition:

$$|1\rangle = \left[\prod_k \frac{1}{|\alpha_k|^{1/2}} e^{-\frac{\beta_k}{2\alpha_k} \hat{a}_k^+ \hat{a}_{-k}^+} \right] |1\rangle \quad (\text{T})$$

needed for normalization

□ Proof: Exercise.

Hint: Use $a e^{2a^+} = e^{2a^+} a + 2e^{2a^+} \dots$

Interpretation of (T):

□ Assume, e.g., that $|1\rangle$ and $|1\rangle$ are those Hilbert space vectors which happen to be the vacuum state vectors at the times η_1 and η_2 respectively.

(We will soon explore how to identify the vacuum state at any given time)

□ Assume at time η_1 , the system's state $|1\rangle$ is the vacuum state (in the sense of no particle state). Then it is convenient to choose the mode functions $\{v_k\}$ so that:

$$|1\rangle = |1\rangle$$

At a later time, η_2 , the system is still in this state but the vacuum is then some other vector $|1\rangle$, which obeys $\hat{a}_k |1\rangle = 0$ with mode functions $\tilde{v}_k(t)$.

□ Thus, from (T) we see that the system's state, $|1\rangle$, is at η_2 a state with many particles.

□ Note: the particles have been created in $k, -k$ pairs.

□ Intuitively: The expansion rips virtual particle + anti-particle pairs apart.