

# Semiclassical Theory of Light Matter Interactions

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## Question (“Semi” in semiclassical?)

- Quantize atom (*observables*  $\leftrightarrow$  *operators*)
- Treat light classically ( $E = E \cos(\nu t)$   $\leftarrow$  *scalar field*)

# Atom-Field Hamiltonian - Obtaining a Hamiltonian

- Start with free-electron Hamiltonian,  $\mathcal{H}_0 = \frac{p^2}{2m}$
- Ask for local phase invariance:

$$\psi(\mathbf{r}, t) \text{ solution} \implies \psi(\mathbf{r}, t)e^{i\theta(\mathbf{r}, t)} \text{ solution}$$

- Introduce the following Hamiltonian:

$$\mathcal{H} = \frac{1}{2m}(\mathbf{p} - i\frac{e}{c}\mathbf{A}(\mathbf{x}, t))^2 + eU(\mathbf{r}, t)$$

with  $(U, \mathbf{A})$  having the following transformation rules:

$$\mathbf{A}(\mathbf{r}, t) \rightarrow \mathbf{A}(\mathbf{r}, t) + \frac{\hbar}{e}\nabla\theta(\mathbf{r}, t)$$

$$U(\mathbf{r}, t) \rightarrow U(\mathbf{r}, t) - \frac{\hbar}{e}\frac{\partial\theta}{\partial t}(\mathbf{r}, t)$$

- Gauge-invariant, hence physical, quantities:

$$\mathbf{E} = -\nabla U - \frac{\partial\mathbf{A}}{\partial t}$$

$$\mathbf{B} = \nabla \times \mathbf{A}$$

# Atom-Field Hamiltonian - Obtaining a Hamiltonian

- Local phase invariance:

$$\mathcal{H} = \frac{1}{2m}(\mathbf{p} - i\frac{e}{c}\mathbf{A}(\mathbf{r}, t))^2 + eU(\mathbf{r}, t) + V(\mathbf{r})$$

- Transform  $\mathcal{H}$  gauge transformation:

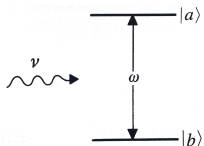
$$\psi(\mathbf{r}, t) \mapsto \exp\left(-\frac{ie}{\hbar}\mathbf{A}(\mathbf{r}_0, t) \cdot \mathbf{r}\right) \psi(\mathbf{r}, t)$$

$$\mathcal{H} = \left[\frac{\mathbf{p}^2}{2m} + V(\mathbf{r})\right] - e\mathbf{r} \cdot \mathbf{E}$$

# Rabi Oscillations

Consider two-level atom in a monochromatic em-field:

$$|\psi(t)\rangle = C_a(t) |a\rangle + C_b(t) |b\rangle$$
$$H = H_0 - e\hat{\mathbf{r}} \cdot \mathbf{E}(\mathbf{r}_0, t)$$



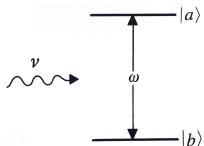
$$H_0 = \hbar\omega_a |a\rangle \langle a| + \hbar\omega_b |b\rangle \langle b|$$
$$H_1 = -e\hat{\mathbf{x}}\mathcal{E} \cos(\nu t)$$
$$= -e\mathcal{E} \cos(\nu t)(\langle a|\hat{\mathbf{x}}|b\rangle + \langle b|\hat{\mathbf{x}}|a\rangle)$$
$$= -\mathcal{E} \cos(\nu t)(\wp_{ab} + \wp_{ba})$$

$\mathcal{E}$  is the amplitude of the field - complex valued.

$\wp_{ab} = \langle a|e\hat{\mathbf{x}}|b\rangle$  is the dipole moment.

# Rabi Oscillations

$$\mathcal{H} = \begin{pmatrix} \hbar\omega_a & \wp_{ab}\mathcal{E} \cos(\nu t) \\ \wp_{ba}\mathcal{E} \cos(\nu t) & \hbar\omega_b \end{pmatrix}$$



Schrodinger equation becomes: ( $\Omega_R = |\wp_{ba}|\mathcal{E}/\hbar$ ,  $\wp_{ba} = |\wp_{ba}|e^{i\phi}$ )

$$\dot{C}_a = -i\omega_a C_a + i\Omega_R e^{-i\phi} \cos(\nu t) C_b$$

$$\dot{C}_b = -i\omega_b C_b + i\Omega_R e^{i\phi} \cos(\nu t) C_a$$

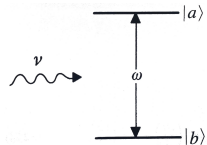
- Transform  $c_\mu := C_\mu e^{i\omega_\mu t}$  for  $\mu = a, b$
- Ignore terms  $\propto e^{\pm i(\omega+\nu)t}$  ← Rotating wave approximation (RWA)

$$\dot{c}_a = i\frac{\Omega_R}{2} e^{-i\phi} c_b e^{i(\omega-\nu)t}$$

$$\dot{c}_b = i\frac{\Omega_R}{2} e^{i\phi} c_a e^{i(\omega-\nu)t}$$

# Rabi Oscillations - Solutions/Special cases

$$\mathcal{H} = \begin{pmatrix} \hbar\omega_a & \wp_{ab}\mathcal{E} \cos(\nu t) \\ \wp_{ba}\mathcal{E} \cos(\nu t) & \hbar\omega_b \end{pmatrix}$$



Solutions:

$$c_a(t) = \left( a_1 e^{i\Omega t/2} + a_2 e^{-i\Omega t/2} \right) e^{i\Delta \cdot t/2}$$

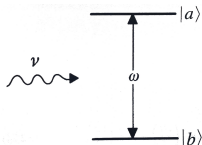
$$c_b(t) = \left( b_1 e^{i\Omega t/2} + b_2 e^{-i\Omega t/2} \right) e^{i\Delta \cdot t/2}$$

where  $\Omega = \sqrt{\Omega_R^2 + (\omega - \nu)^2}$ .



# Rabi Oscillations - Solutions/Special cases

$$\mathcal{H} = \begin{pmatrix} \hbar\omega_a & \wp_{ab}\mathcal{E} \cos(\nu t) \\ \wp_{ba}\mathcal{E} \cos(\nu t) & \hbar\omega_b \end{pmatrix}$$

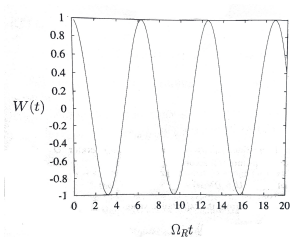


Special Case: detuning  $\Delta = \omega - \nu = 0$  induce Rabi oscillations.

$$c_a(t) = \left[ c_a(0) \cos\left(\frac{\Omega_R t}{2}\right) + i c_b(0) \sin\left(\frac{\Omega_R t}{2}\right) \right]$$
$$c_b(t) = \left[ c_b(0) \cos\left(\frac{\Omega_R t}{2}\right) + i c_a(0) \sin\left(\frac{\Omega_R t}{2}\right) \right]$$

Inversion:

$$W(t) = |c_a(t)|^2 - |c_b(t)|^2$$



Polarization - expectation of dipole-moment operator:

$$\begin{aligned} P(t) &= e \langle \psi(t) | \hat{r} | \psi(t) \rangle \\ &= c_a^* c_b \rho_{ab} e^{i\omega t} + c.c. \end{aligned}$$

# Optical Bloch Equations

Density matrices,  $\rho$ , characterized by

- $\text{Tr}(\rho) = 1$ ,
- $\rho = \rho^\dagger$ ,
- $\langle \psi | \rho | \psi \rangle \geq 0$  for all  $|\psi\rangle$ .

Von-Neumann equation of motion:

$$i\hbar\dot{\rho} = [\mathcal{H}, \rho]$$

In two level atom:

$$\rho = \begin{pmatrix} \rho_{aa} & \rho_{ab} \\ \rho_{ba} & \rho_{bb} \end{pmatrix}$$

- Diagonal elements correspond to populations of atoms
- Off-diagonal elements, called *coherences*, correspond to response at driving frequency

# Optical Bloch Equations - with noise

\*\*Spontaneous emission is taken into account by introducing a dissipative matrix,  $\Gamma$ , with matrix element:

$$\langle n|\Gamma|m\rangle = \gamma_n\delta_{nm}$$

The equation of motion becomes:

$$i\hbar\dot{\rho} = [\mathcal{H}, \rho] - \frac{i\hbar}{2}(\Gamma\rho + \rho\Gamma)$$

In component form (Optical Bloch Equations):

$$\dot{\rho}_{aa} = -\gamma_1\rho_{aa} + \frac{i}{\hbar}[\wp_{ab}E\rho_{ba} - cc]$$

$$\dot{\rho}_{bb} = -\gamma_2\rho_{bb} - \frac{i}{\hbar}[\wp_{ab}E\rho_{ba} - cc]$$

$$\dot{\rho}_{ab} = -(i\omega + \gamma_{ab})\rho_{ab} - \frac{i}{\hbar}\wp_{ab}E(\rho_{aa} - \rho_{bb})$$

Formulation allows for more effects!

Slowly working towards a theory of the laser...

- Atoms pumped to excited state
- Interact with EM-field

# Maxwell-Schrodinger – Population Matrix

- Atoms will not interact with field until pumped to excited state

$$\rho(z, t, t_0) - \text{pumped at } t_0$$

- $r_a(z, t_0)$  – pumping rate (atoms per second per volume)
- Population matrix – the density matrix of all the atoms in a cavity

$$\rho(z, t) = \int_{-\infty}^t dt_0 r_a(z, t_0) \rho(z, t, t_0)$$

# Maxwell-Schrodinger – Population Matrix

## Macroscopic Polarization

$$\begin{aligned}\rho(z, t) &= \int_{-\infty}^t dt_0 r_a(z, t_0) \rho(z, t, t_0) \\ &= \sum_{\alpha, \beta} \int_{-\infty}^t dt_0 r_a(z, t_0) \rho_{\alpha, \beta}(z, t, t_0) |\alpha\rangle \langle \beta|\end{aligned}$$

Macroscopic polarization - crucial to tie in with Maxwell's equations:

$$P(z, t) = \text{Tr}(\hat{\rho}\rho(z, t)) = \wp_{ab}\rho_{ab}(z, t) + cc$$

# Maxwell-Schrodinger – Population Matrix

## Equation of Motion

Differentiating,

$$\rho(z, t) = \sum_{\alpha, \beta} \int_{-\infty}^t dt_0 r_a(z, t_0) \rho_{\alpha, \beta}(z, t, t_0) |\alpha\rangle \langle \beta|$$

We get the “Schrodinger” part of Maxwell-Schrodinger:

$$\dot{\rho}_{aa} = \lambda_a - \gamma_a \rho_{aa} + \frac{i}{\hbar} [\wp_{ab} E \rho_{ba} - cc]$$

$$\dot{\rho}_{bb} = \lambda_b - \gamma_b \rho_{bb} - \frac{i}{\hbar} [\wp_{ab} E \rho_{ba} - cc]$$

$$\dot{\rho}_{ab} = -(i\omega + \gamma) \rho_{ab} - \frac{i}{\hbar} \wp_{ab} E (\rho_{aa} - \rho_{bb})$$

where  $\lambda_a = r_a \rho_{aa}^{(0)}$ ,  $\lambda_b = r_b \rho_{bb}^{(0)}$ .



Maxwell's Equations:

$$\begin{aligned}\nabla \cdot \mathbf{D} &= 0 & \nabla \times \mathbf{E} &= -\frac{\partial \mathbf{B}}{\partial t} \\ \nabla \cdot \mathbf{B} &= 0 & \nabla \times \mathbf{H} &= \mathbf{J} + \frac{\partial \mathbf{D}}{\partial t},\end{aligned}$$

where

$$\mathbf{D} = \epsilon_0 \mathbf{E} + \mathbf{P}, \quad \mathbf{B} = \mu_0 \mathbf{H}, \quad \mathbf{J} = \sigma \mathbf{E}$$

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Then:

$$\nabla \times (\nabla \times \mathbf{E}) + \mu_0 \sigma \frac{\partial \mathbf{E}}{\partial t} + \mu_0 \epsilon_0 \frac{\partial^2 \mathbf{E}}{\partial t^2} = -\mu_0 \frac{\partial^2 \mathbf{P}}{\partial t^2}$$

Assume, unidirectional electric field inside cavity,  $\mathbf{E}(\mathbf{r}, t) = E(z, t)\hat{\mathbf{x}}$

$$E(z, t) = \frac{1}{2}\mathcal{E}(z, t)e^{-i[\nu t - kz + \phi(z, t)]} + cc$$

$$P(z, t) = \frac{1}{2}\mathcal{P}(z, t)e^{-i[\nu t - kz + \phi(z, t)]} + cc$$

where  $\mathcal{E}$ ,  $\mathcal{P}$ ,  $\phi$  are all slowly-varying functions.

$$\mathcal{P}(z, t) := 2\wp\rho_{ab}e^{i[\nu t - kz + \phi(z, t)]}$$

Now, we find equations of motion for  $\mathcal{E}$ ,  $\phi$ .

$$-\frac{\partial^2 E}{\partial z^2} + \mu_0 \sigma \frac{\partial E}{\partial t} + \frac{1}{c^2} \frac{\partial^2 E}{\partial t^2} = -\mu_0 \frac{\partial^2 P}{\partial t^2}$$

We arrive at two equations that describe  $\mathcal{E}$ ,  $\phi$ :

$$\begin{aligned} \frac{\partial \mathcal{E}}{\partial z} + \frac{1}{c} \frac{\partial \mathcal{E}}{\partial t} &= -\kappa \mathcal{E} - \frac{1}{2\epsilon_0} k \operatorname{Im} \mathcal{P} \\ \frac{\partial \phi}{\partial z} + \frac{1}{c} \frac{\partial \phi}{\partial t} &= k - \frac{\nu}{c} - \frac{1}{2\epsilon_0} k \mathcal{E}^{-1} \operatorname{Re} \mathcal{P} \end{aligned}$$

Self-consistent:  $E$  driven by population-matrix elements <sup>1</sup>

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<sup>1</sup>PRA 81 033833

# Stimulated emission and absorption

Suppose  $\Delta = \omega - \nu = 0$ ,  $c_a(0) = 1$ ,  $c_b(0) = 0$ ; solution becomes:

$$c_a(t) = \cos\left(\frac{\Omega_R t}{2}\right) \approx 1$$

$$c_b(t) = i \sin\left(\frac{\Omega_R t}{2}\right) \approx i \frac{\Omega_R t}{2}$$

Polarization

$$\mathcal{P} = 2\varphi \rho_{ab} e^{i\nu t} \approx -i\varphi \Omega_R t$$

Maxwell's Equation gives:

$$\frac{1}{c} \frac{\partial \mathcal{E}}{\partial t} = \frac{k}{2\epsilon_0} \frac{\varphi^2}{\hbar} \mathcal{E} t^2 \implies \Delta \mathcal{E} \approx \frac{ck}{2\epsilon_0} \frac{\varphi^2}{\hbar} \mathcal{E} t^2$$

If  $c_a(0) = 0$ ,  $c_b(0) = 1$  then  $\Delta \mathcal{E} \approx -\frac{ck}{2\epsilon_0} \frac{\varphi^2}{\hbar} \mathcal{E} t^2$

# Appendix

# Maxwell-Schrodinger

Maxwell's Part – almost full derivation

$$\begin{aligned}\nabla \times (\nabla \times \mathbf{E}) + \mu_0 \sigma \frac{\partial \mathbf{E}}{\partial t} + \mu_0 \epsilon_0 \frac{\partial^2 \mathbf{E}}{\partial t^2} &= -\mu_0 \frac{\partial^2 \mathbf{P}}{\partial t^2} \\ \implies -\frac{\partial^2 \mathbf{E}}{\partial z^2} + \mu_0 \sigma \frac{\partial \mathbf{E}}{\partial t} + \frac{1}{c^2} \frac{\partial^2 \mathbf{E}}{\partial t^2} &= -\mu_0 \frac{\partial^2 \mathbf{P}}{\partial t^2} \\ \implies \left( \frac{\partial}{\partial z} + \frac{1}{c} \frac{\partial}{\partial t} \right) \underbrace{\left( -\frac{\partial}{\partial z} + \frac{1}{c} \frac{\partial}{\partial t} \right)}_{\simeq -2ik} \mathbf{E} &= -\mu_0 \sigma \frac{\partial \mathbf{E}}{\partial t} - \mu_0 \frac{\partial^2 \mathbf{P}}{\partial t^2}\end{aligned}$$

We arrive at two equations that describe  $\mathcal{E}, \phi$ :

$$\begin{aligned}\frac{\partial \mathcal{E}}{\partial z} + \frac{1}{c} \frac{\partial \mathcal{E}}{\partial t} &= -\kappa \mathcal{E} - \frac{1}{2\epsilon_0} k \operatorname{Im} \mathcal{P} \\ \frac{\partial \phi}{\partial z} + \frac{1}{c} \frac{\partial \phi}{\partial t} &= k - \frac{\nu}{c} - \frac{1}{2\epsilon_0} k \mathcal{E}^{-1} \operatorname{Re} \mathcal{P}\end{aligned}$$

# Semiclassical laser theory

Quick sketch of §5.5 Scully

- 1 Integrate off-diagonal Schrodinger equation to obtain  $\mathcal{P}(z, t)$

$$\dot{\rho}_{ab} = -(i\omega + \gamma)\rho_{ab} - \frac{i}{\hbar} \wp \mathcal{E}(\rho_{aa} - \rho_{bb})$$

- 2 Substitute into the other Schrodinger equation:

$$\begin{aligned}\dot{\rho}_{aa} &= \lambda_a - \gamma_a \rho_{aa} - R(\rho_{aa} - \rho_{bb}) \\ \dot{\rho}_{bb} &= \lambda_b - \gamma_b \rho_{bb} + R(\rho_{aa} - \rho_{bb})\end{aligned}$$

where  $R = \frac{1}{2} \left( \frac{\wp \mathcal{E}}{\hbar} \right)^2 \frac{\gamma}{\gamma^2 + (\omega - \nu)^2}$ .

- 3 Solve for  $\rho_{aa} - \rho_{bb}$  in steady state, obtain  $\mathcal{P}$
- 4 Introduce dimensionless intensity (“photon number”)

$$n = \frac{\epsilon_0 \mathcal{E}^2 V}{2\hbar\nu}$$

- 5 Rewrite Maxwell's equations

# Time-Dependent Perturbation

Let  $H = H_0 + \lambda V(t)$ , so that  $H_0$  is time-independent and  $\lambda V(t) \ll 1$ .

In the interaction picture, time dependence only due to interaction:

$$|\psi(t)\rangle_D = U_0 |\psi(t)\rangle$$

$$A(t)_D = U_0 A(t) U_0^\dagger$$

we have the following evolution:

$$i \frac{d}{dt} |\psi(t)\rangle_D = \lambda \hat{V}_D |\psi(t)\rangle_D$$



# Time-Dependent Perturbation

$$H = H_0 + \lambda V(t)$$

$$|\psi(t)\rangle_D = U_0 |\psi(t)\rangle$$

$$A(t)_D = U_0 A(t) U_0^\dagger$$

$$i \frac{d}{dt} |\psi(t)\rangle_D = \lambda \hat{V}_D |\psi(t)\rangle_D$$

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Take  $U_D(t, t_0)$  so that  $|\psi(t)\rangle_D = U_D(t, t_0) |\psi(t_0)\rangle_D$  then:

$$\frac{d}{dt} U_D(t, t_0) = -i\lambda V_D(t) U_D(t, t_0)$$

Integrating,

$$U_D(t, t_0) = \mathbb{1} - i\lambda \int_{t_0}^t d\tau V_D(\tau) U_D(\tau, t_0)$$

$$\begin{aligned} U_D(t, t_0) &= \mathbb{1} - i\lambda \int_{t_0}^t d\tau_1 V_D(\tau_1) \left( \mathbb{1} - i\lambda \int_{t_0}^{\tau_1} d\tau_2 V_D(\tau_2) (\dots) \right) \\ &= \mathbb{1} + U_D^{(1)}(t, t_0) + U_D^{(2)}(t, t_0) + \dots \end{aligned}$$